

Radioactivity

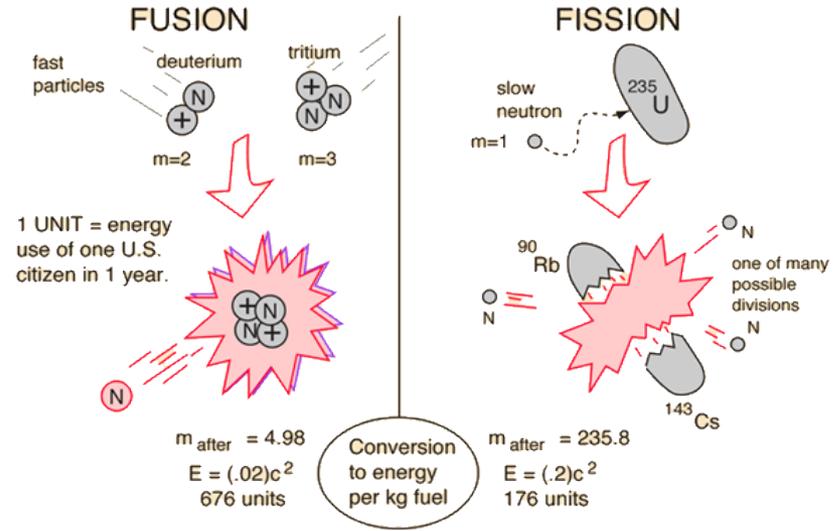
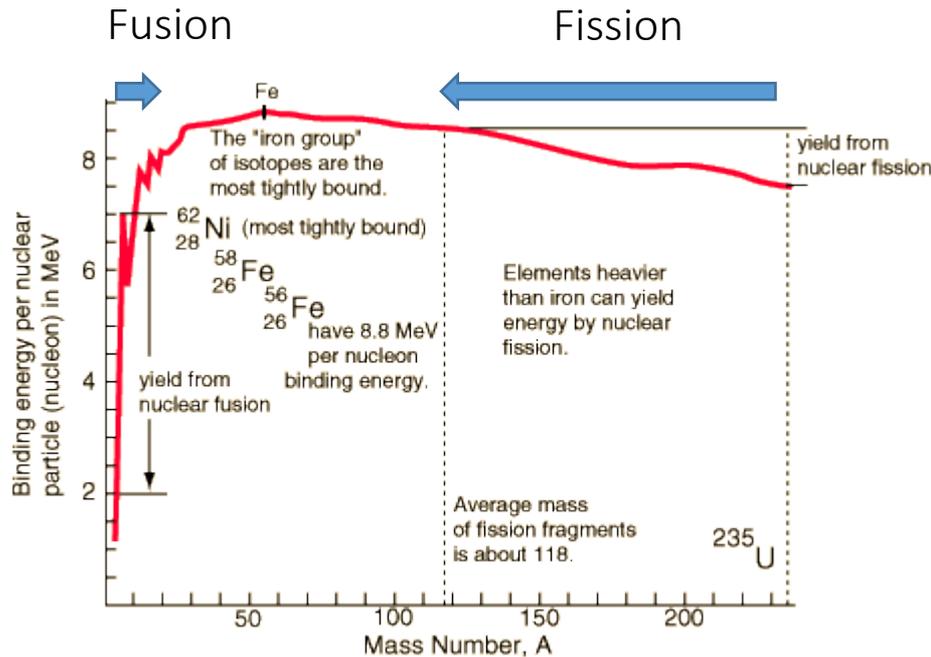
Lecture 20

Radioactivity and Nuclear Energy

What are the issues with Nuclear Energy?

- Nuclear energy is the efficient energy source, it transforms matter into energy according to $E=m\cdot c^2$. In energy generation efficiency nuclear energy is superior to all other sources!
- It has a price, multiple issues of technical and emotional nature!
- Fusion versus Fission
- The issues of fission technologies:
 - fuel production through mining
 - fuel efficiency by different fission modes,
 - fuel burning control by neutron moderation and absorption
 - fuel poisoning through fission products
 - what to do with the fission products?
- The handling of nuclear waste!

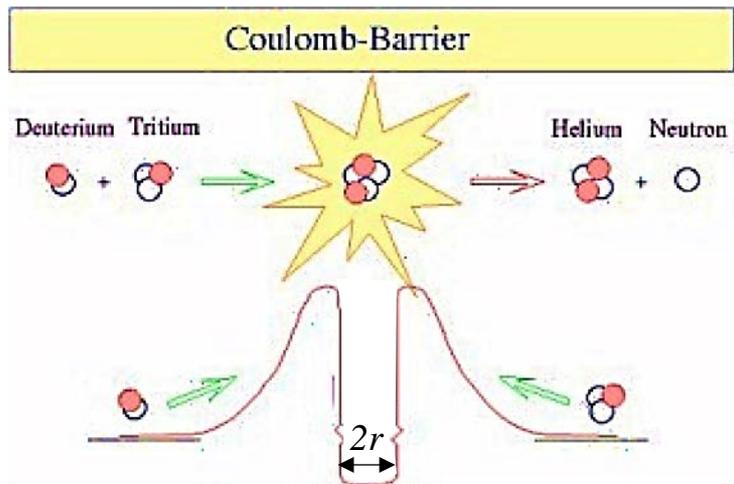
Nuclear Energy Production



The energy release per atomic mass unit is 0.7MeV for fission and 6.2MeV for fusion, fusion is the more effective nuclear reaction!

Fusion occurs by the nuclear reaction between two light hydrogen isotopes, such as d+d ($^2\text{H}+^2\text{H}$), or d+t ($^2\text{H}+^3\text{H}$), while fission is the neutron induced splitting of ^{235}U or ^{239}Pu into two lower mass isotopes between mass 100 and mass 130.

Challenges in fusion



$$r = r_0 \cdot A^{1/3}$$

$$r_0 = 1.25 \text{ fm} = 1.25 \cdot 10^{-13} \text{ cm}$$

$$\begin{aligned}
 U(r) &= \frac{Z_1 Z_2 e^2}{4\pi\epsilon_0 r} = \frac{(1)(1)}{4(3.14)(8.85 \times 10^{-12} \text{ C}^2 \text{ N}^{-1} \text{ m}^{-2})} \frac{(1.6 \times 10^{-19} \text{ C})^2}{1 \times 10^{-15} \text{ m}} \\
 &= (9.01 \times 10^9)(2.56 \times 10^{-23}) \text{ Nm} \\
 &= 2.31 \times 10^{-13} \text{ J} \cdot \frac{1 \text{ eV}}{1.6 \times 10^{-19} \text{ J}} \\
 &= 1.44 \text{ MeV} \\
 1 \text{ J} &= 6.242 \cdot 10^{12} \text{ MeV}
 \end{aligned}$$

Typical energy of particles in sun: $E=k \cdot T$, with $T=15 \text{ MK}$ and $k=1.38064852 \cdot 10^{-23} \text{ m}^2 \text{ kg s}^{-2} \text{ K}^{-1}$
 $=4.47 \cdot 10^{-18} \text{ J} = 2.8 \cdot 10^{-5} \text{ MeV} = 28 \text{ eV}$: far below the energy of the Coulomb barrier.

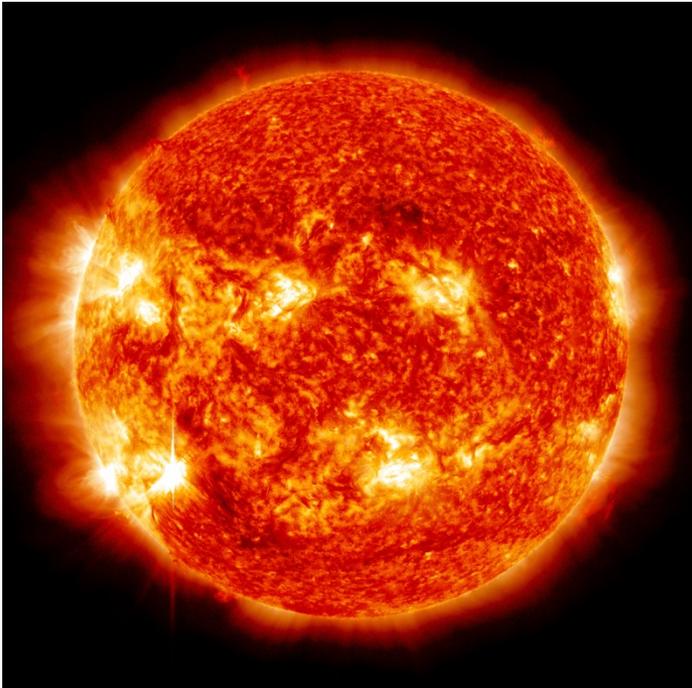
Because of quantum mechanical wave behavior of particles, they can tunnel through Coulomb barrier with a certain probability, which determines timescale of slow stellar burning.

Typical energy of particles on Earth: $E=k \cdot T$, with $T=300 \text{ K}$ and $k=1.38064852 \cdot 10^{-23} \text{ m}^2 \text{ kg s}^{-2} \text{ K}^{-1}$
 $=4.14 \cdot 10^{-21} \text{ J} = 2.6 \cdot 10^{-8} \text{ MeV} = 26 \text{ meV}$: even further below the energy of the Coulomb barrier.

Earth temperatures are far too low for spontaneous fusion, artificial hot plasmas are needed for bringing the particles together at a reasonable rate for sufficient energy generation!

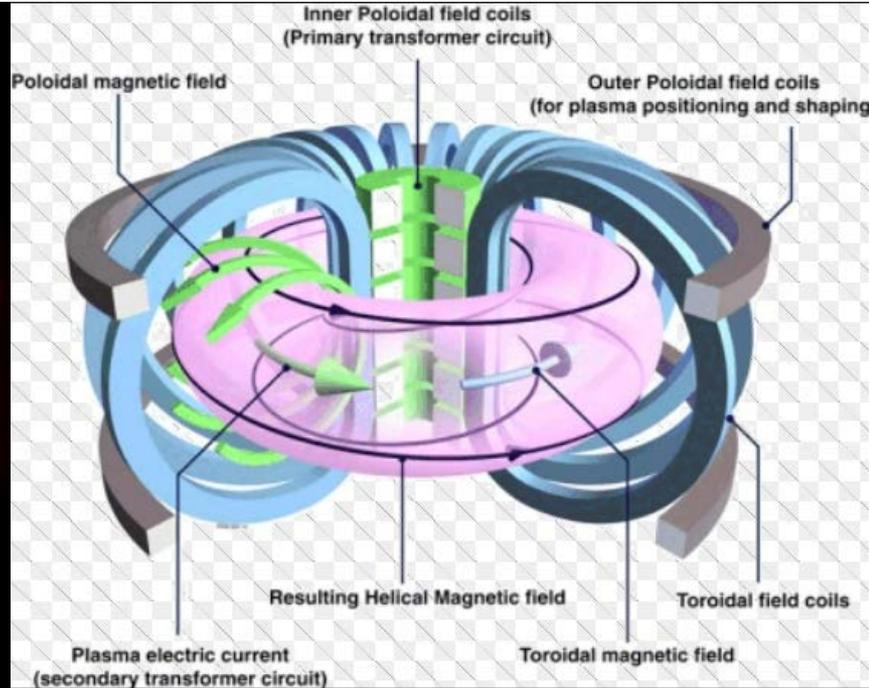
Confinement of hot Plasmas

Gravitational
confinement



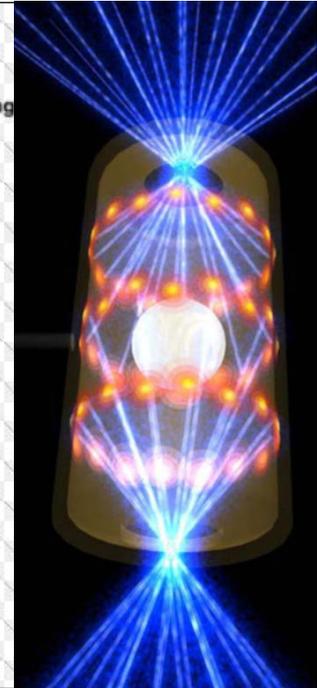
Nature's approach

Magnetic
confinement



human approach

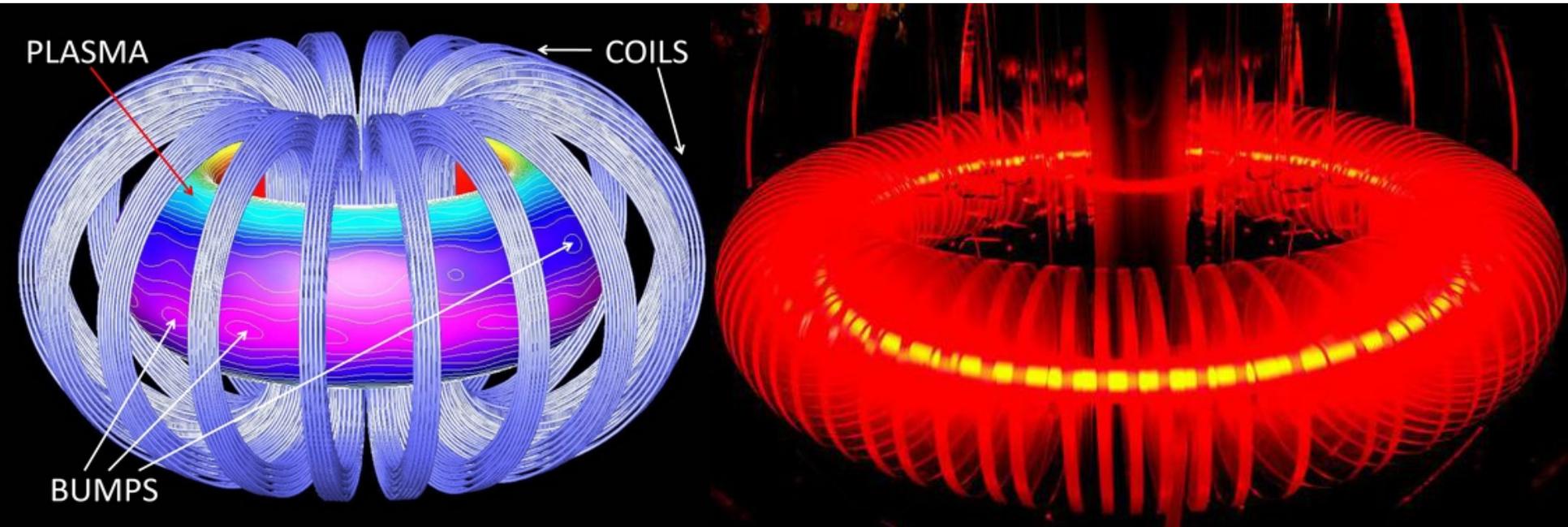
Inertial
confinement



Plasma fusion

through magnetic confinement

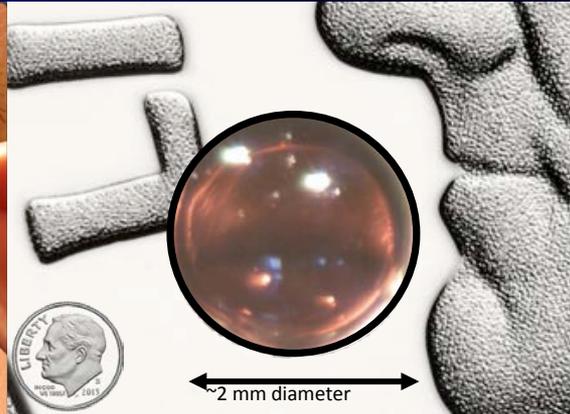
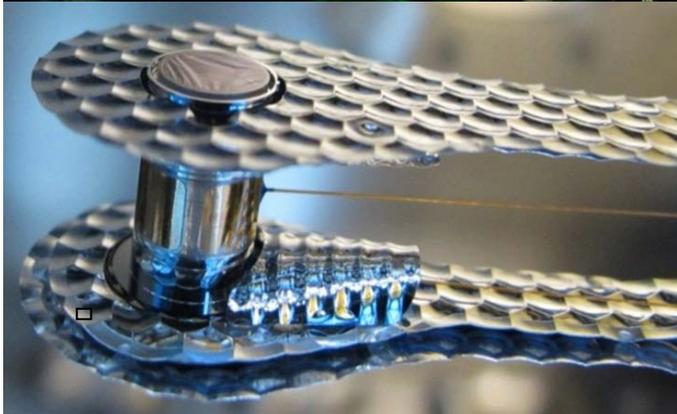
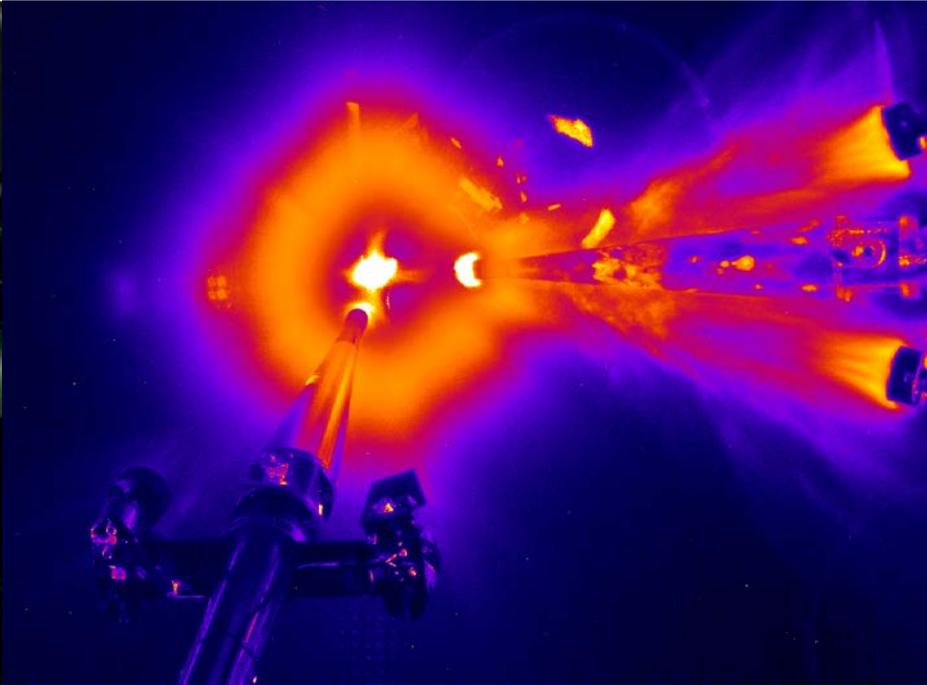
A major advantage of fusion reactors is the small if not negligible amount long-lived radioactive decay products! Light radioactive isotopes are short-lived and produce additional energy through the decay heat.



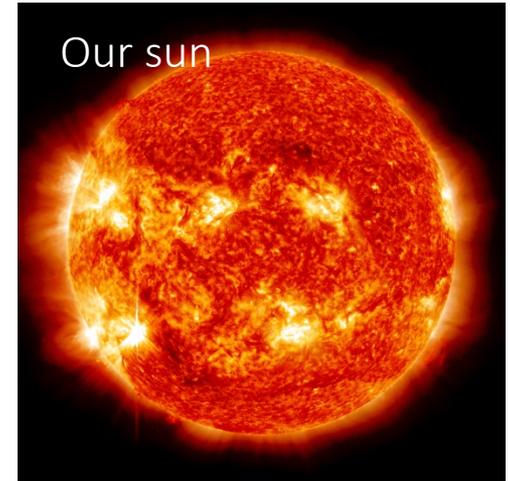
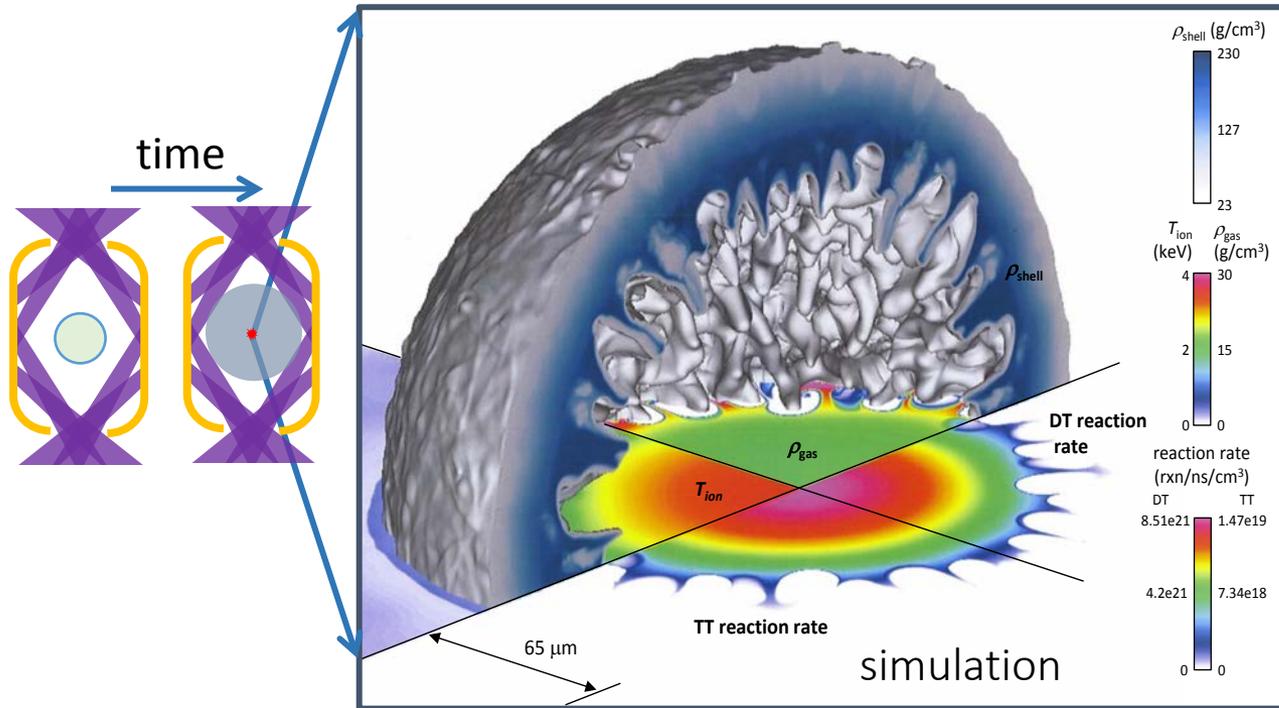
Plasma has to have temperatures of ten times the temperature in the core of the sun to generate the required energy output. The plasma is contained by magnetic fields 10,000 times that of the Earth's field. The shape of the fusion plasma is dictated by the magnetic field generation. These provides enormous technical challenges that have not been achieved yet. The main project towards the goal is the international ITER project in France, but there are still a number of smaller projects in the US, CHINA, and Germany.

Laser Induced Fusion through inertial confinement

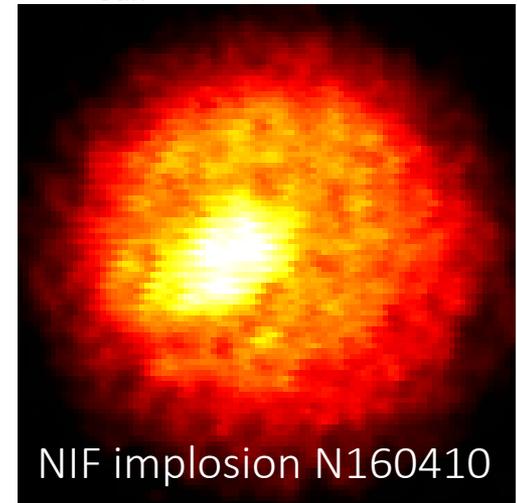
195 Peta-Watt Lasers aiming at one spot generating an implosion of capsule reaching temperatures and densities close or superseding solar values.



NIF shot conditions



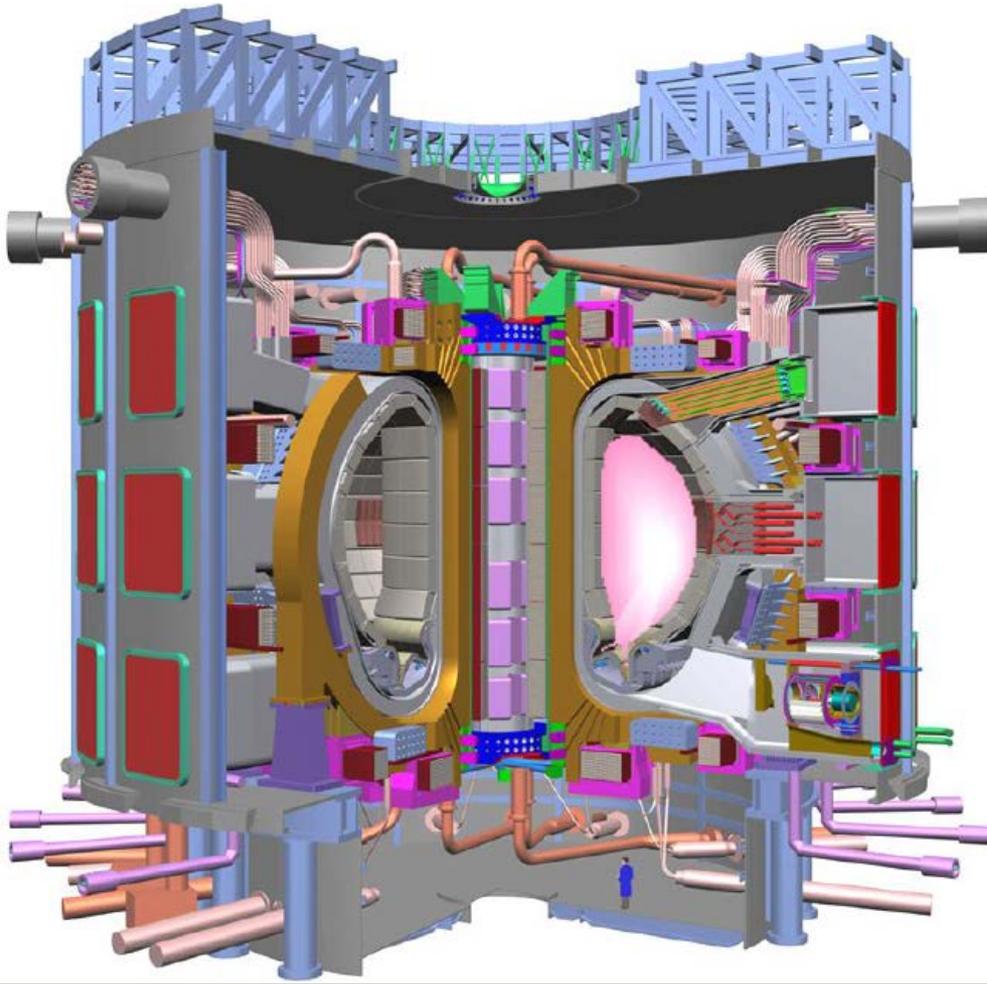
$M_{\text{Sun}} \sim 2 \times 10^{33} \text{g}$



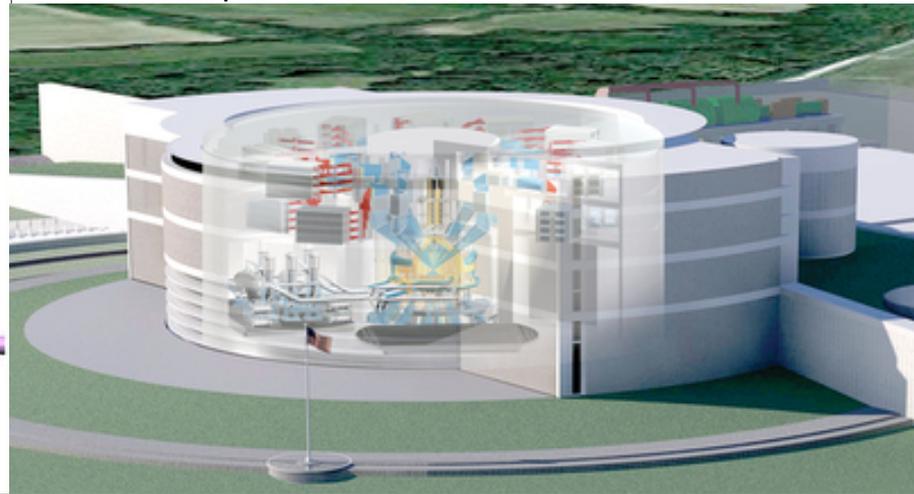
$M_{\text{impl}} \sim 6 \times 10^{-6} \text{g}$

Shot physics is not fully understood yet. Rapid convective processes seems to inhibit the production and release of positive net energy. A problem is the limited frequency or shot-rate, three shots/day, desired rate is 100 shots/second, a three million times improvement is necessary!

Time-line 30-50 years



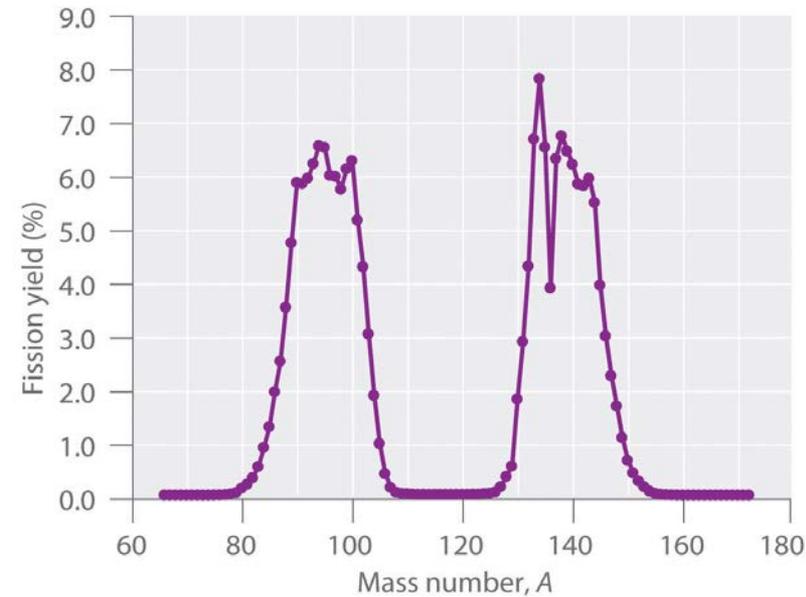
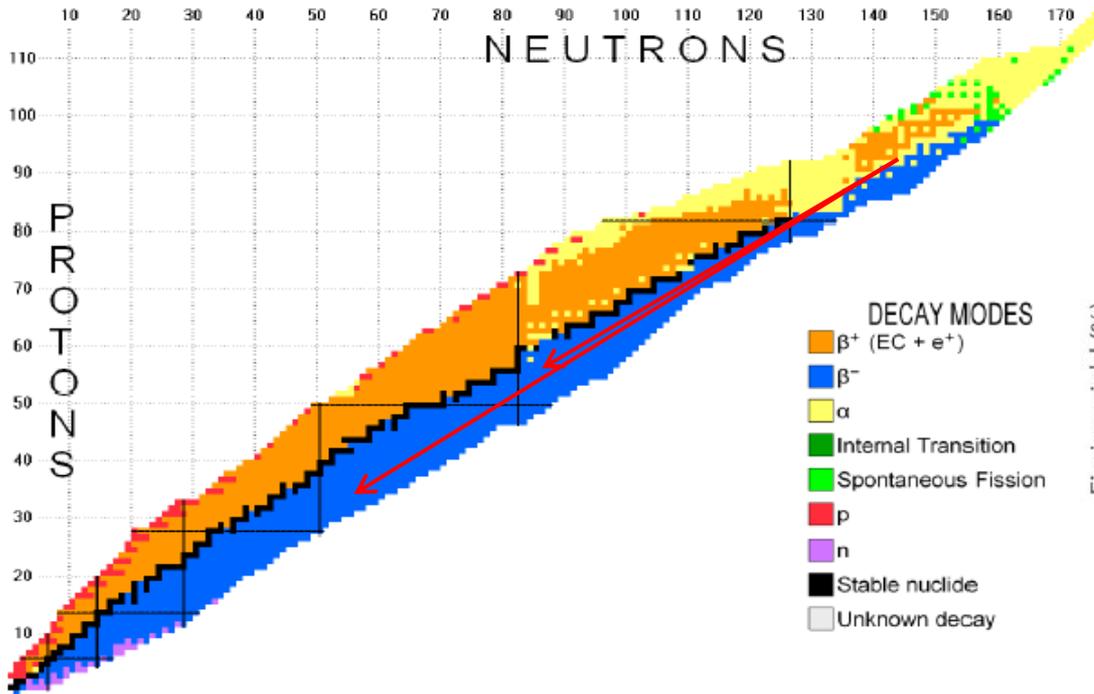
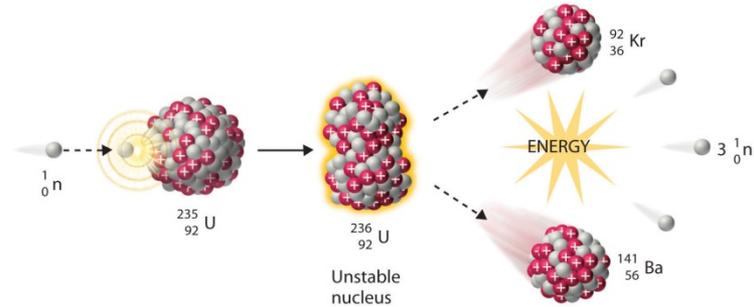
The general statement is that it takes 30-50 years towards reliable energy production by fusion, but the world cannot wait, the energy demand is growing exponentially! Nuclear fission is one of the existing options for bridging the time until a reliable fusion concept is developed!



Present concepts: ITER

LIFE

Challenges in Fission

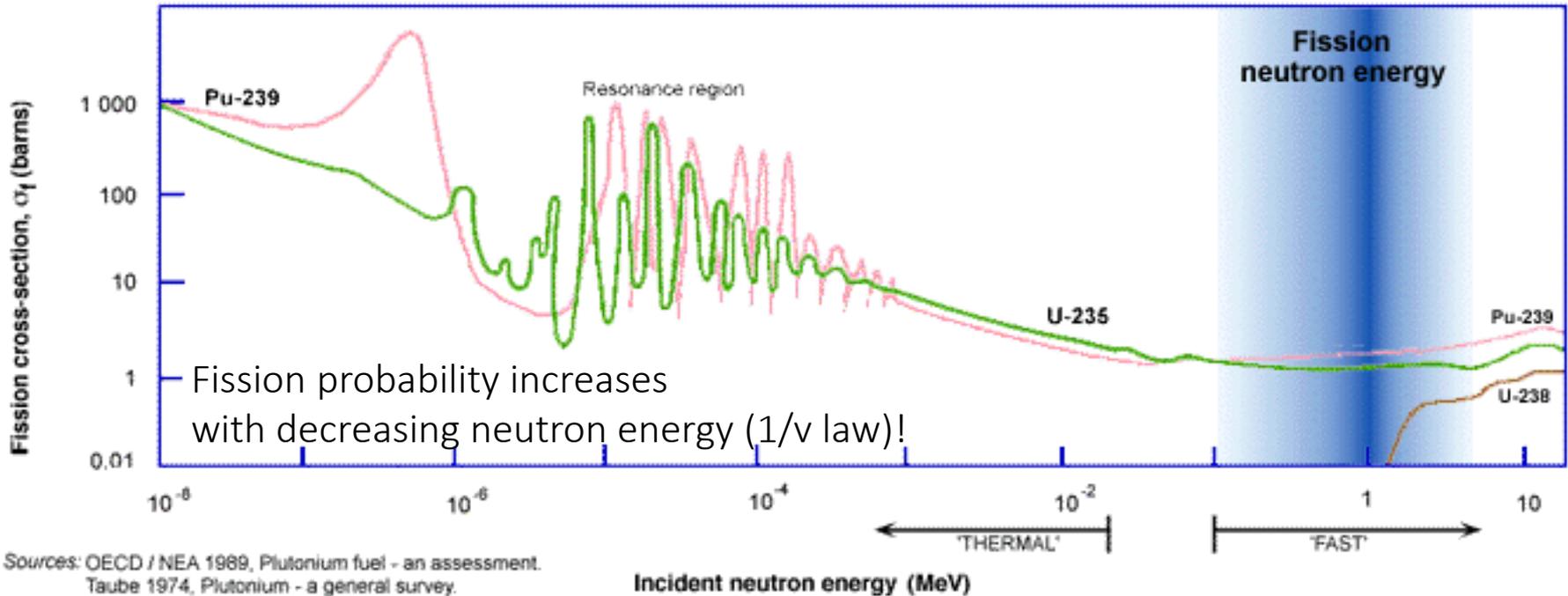


Heavy nuclei can fission by breaking up in two lower mass nuclei. The fission products come in a double bump distribution of pairs of neutron rich radioactive isotopes. Fission produces long-lived radioactivity.

Each fission event produces additional free neutrons that trigger the next fission event in dense fissionable material.

Thermalisation

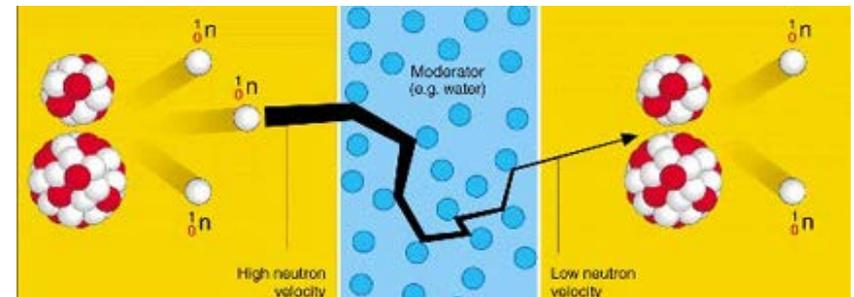
NEUTRON CROSS-SECTIONS FOR FISSION OF URANIUM AND PLUTONIUM



Sources: OECD / NEA 1989, Plutonium fuel - an assessment.
Taube 1974, Plutonium - a general survey.
1 barn = 10^{-28} m², 1 MeV = 1.6×10^{-13} J

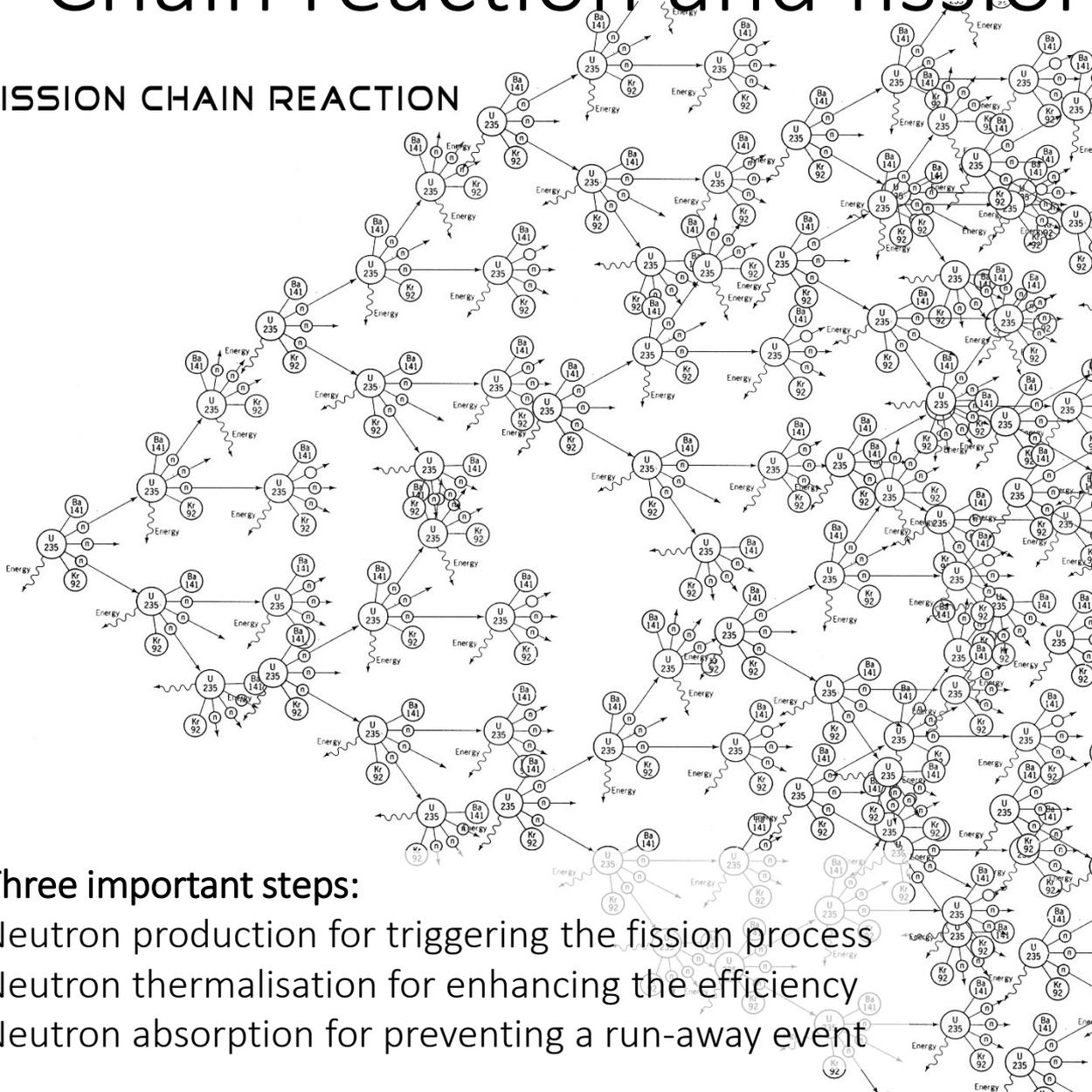
Fission neutrons are emitted with high energy and have to be slowed down (moderated) to be efficiently captured by uranium or plutonium.

The moderator material must be low Z material to pass on a large amount of energy from neutron to scattering particle. The material also must have a low cross section for neutron absorption since the high neutron flux needs to be maintained.

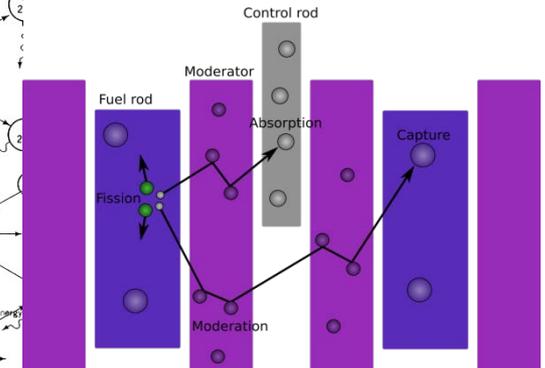


Chain reaction and fission cycles

FISSION CHAIN REACTION



controlled



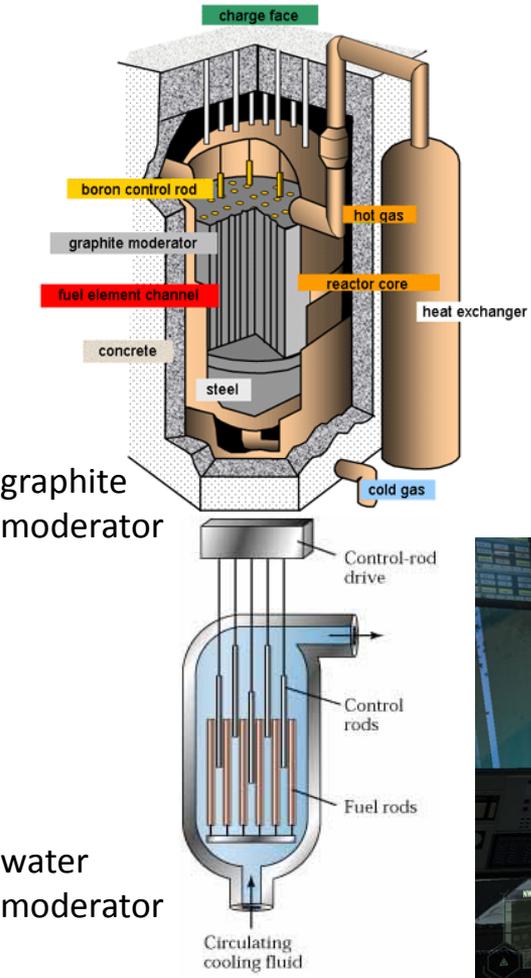
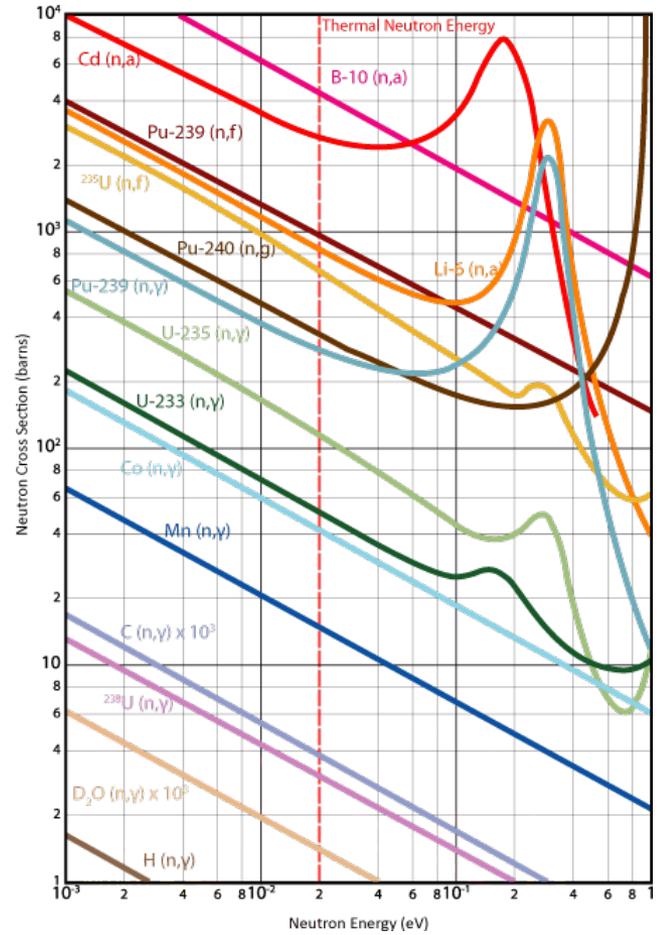
Three important steps:

- Neutron production for triggering the fission process
- Neutron thermalisation for enhancing the efficiency
- Neutron absorption for preventing a run-away event

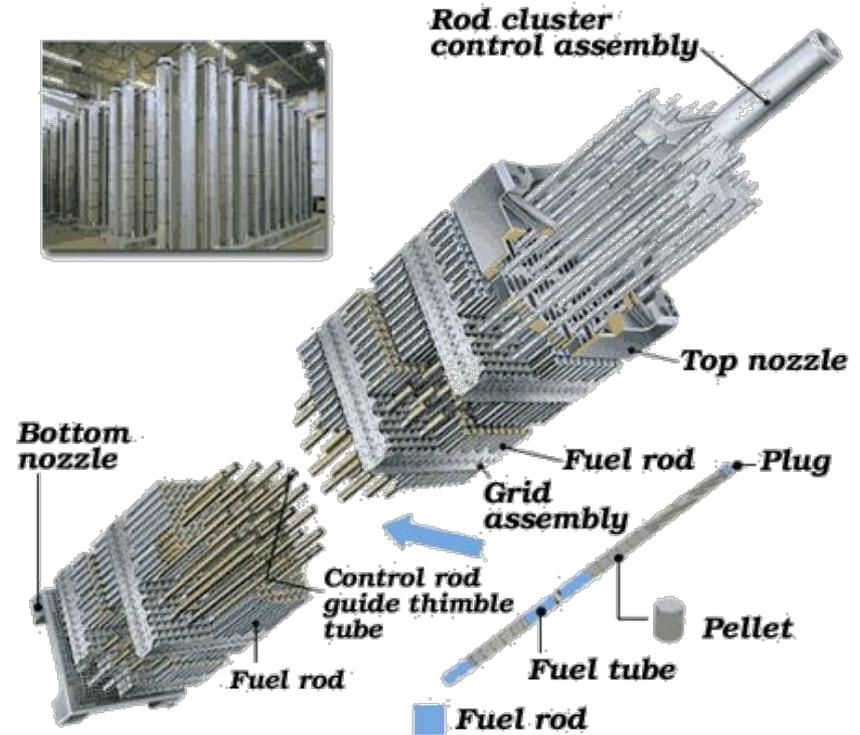
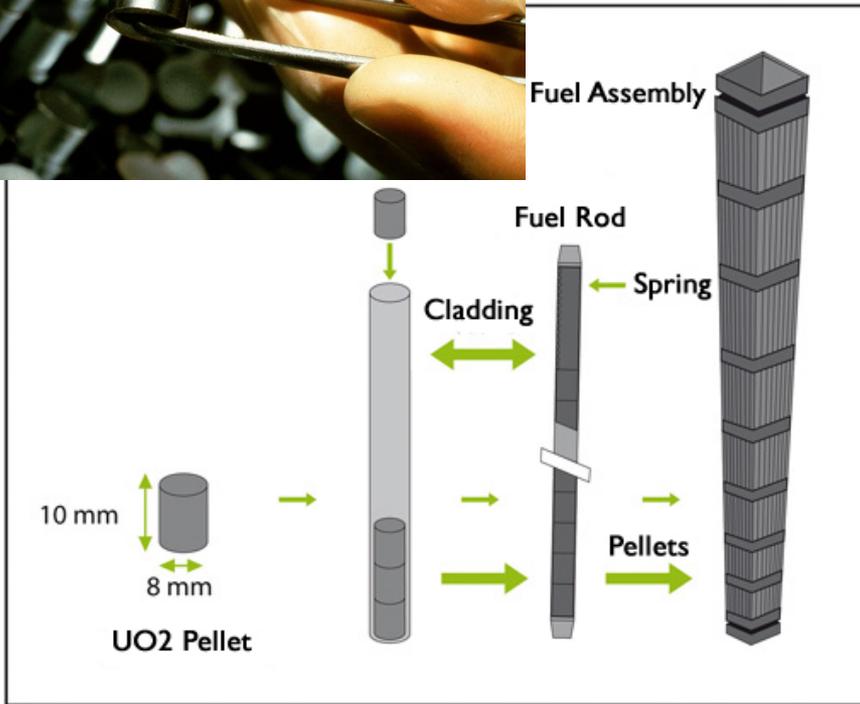
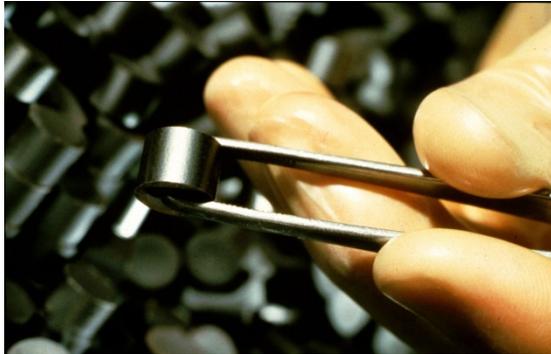
Control rods

The control rods are embedded in a matrix between fuel elements and moderators. Their main mission is to control the neutron flux to prevent the reactor from becoming critical.

The two favorite materials are Cadmium and Boron, because of their high neutron capture cross section. The reaction $^{10}\text{B}(n,\alpha)^7\text{Li}$, produces stable ^7Li that can easily be removed. Neutron absorption on cadmium at thermal energies is driven by $^{113}\text{Cd}(n,\gamma)^{114}\text{Cd}$, forming a stable ^{114}Cd isotope. Subsequent neutron capture on ^{114}C produces short-lived ^{115}Cd .



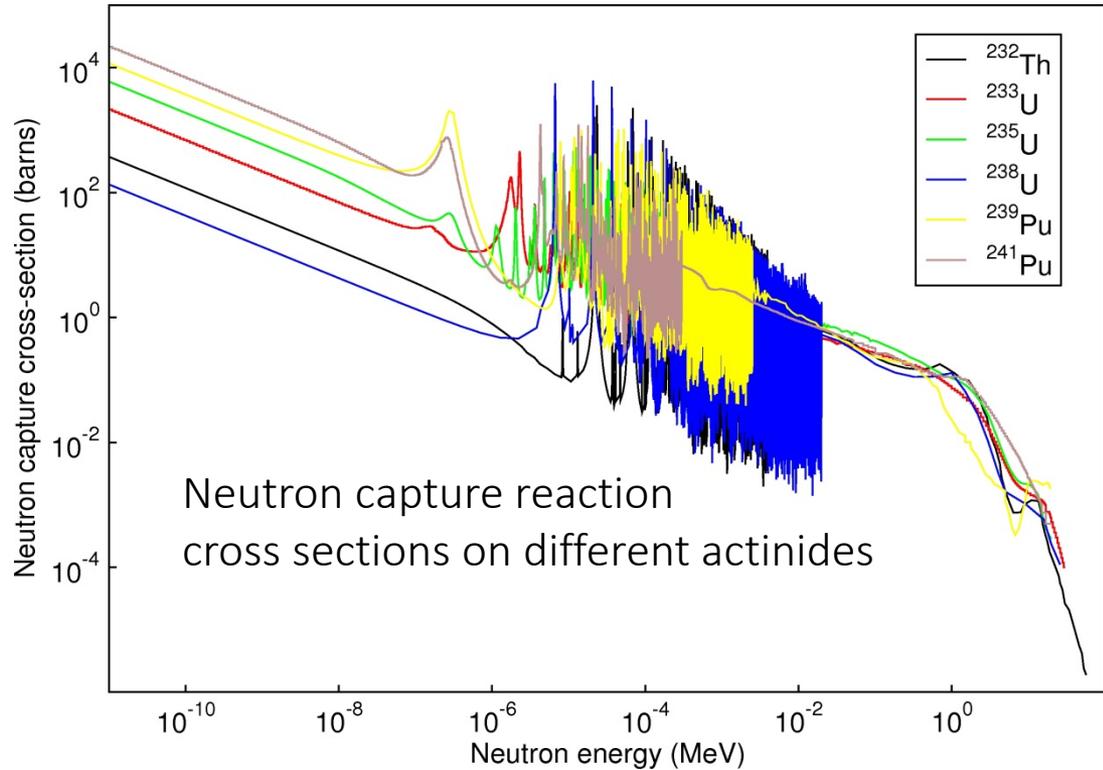
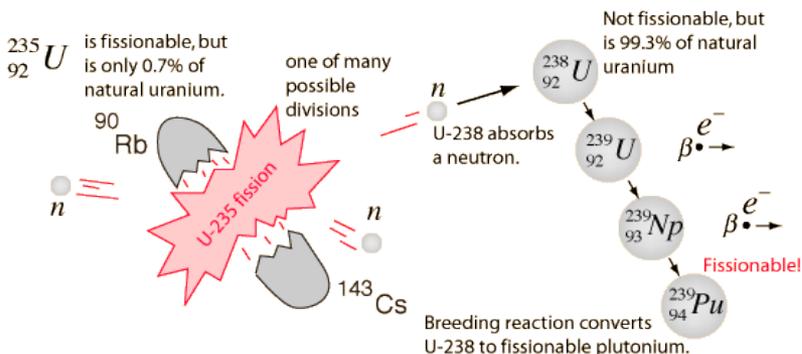
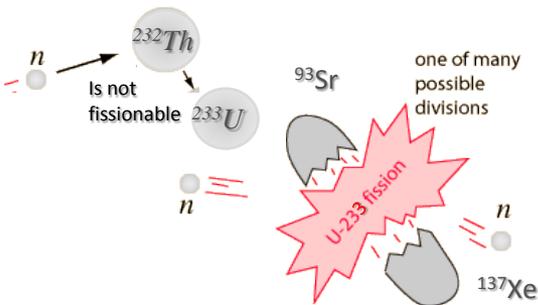
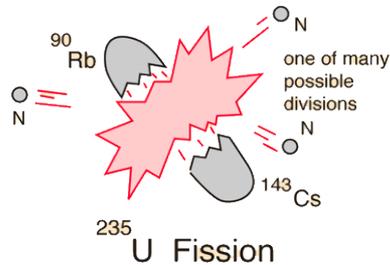
Fuel elements



The assembly of fuel and moderator elements is an engineering problem towards optimized operation, safety, assembly, and disassembly for storage and recycling.

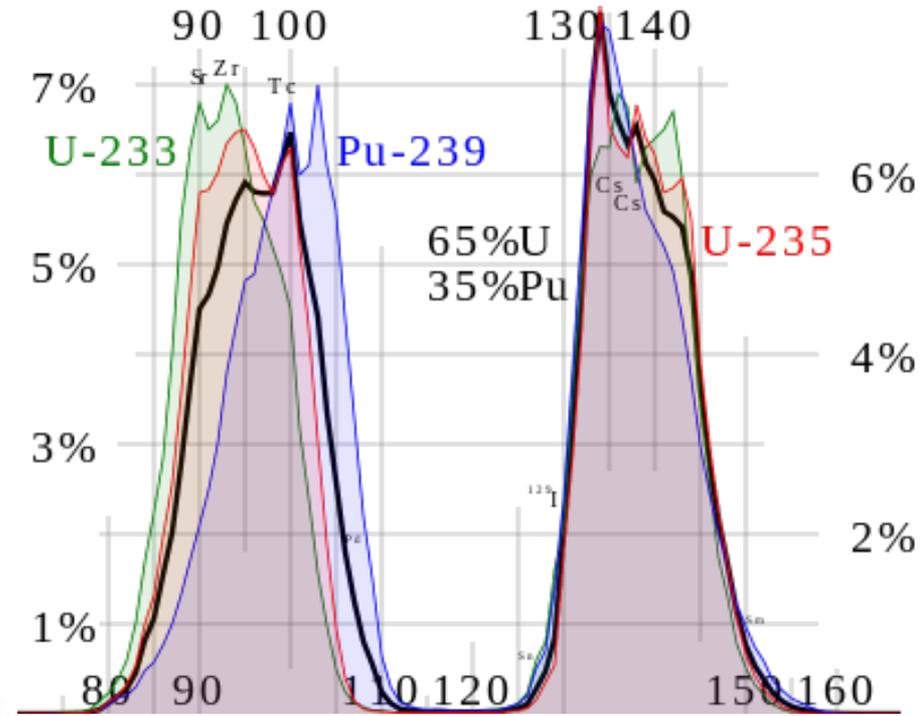
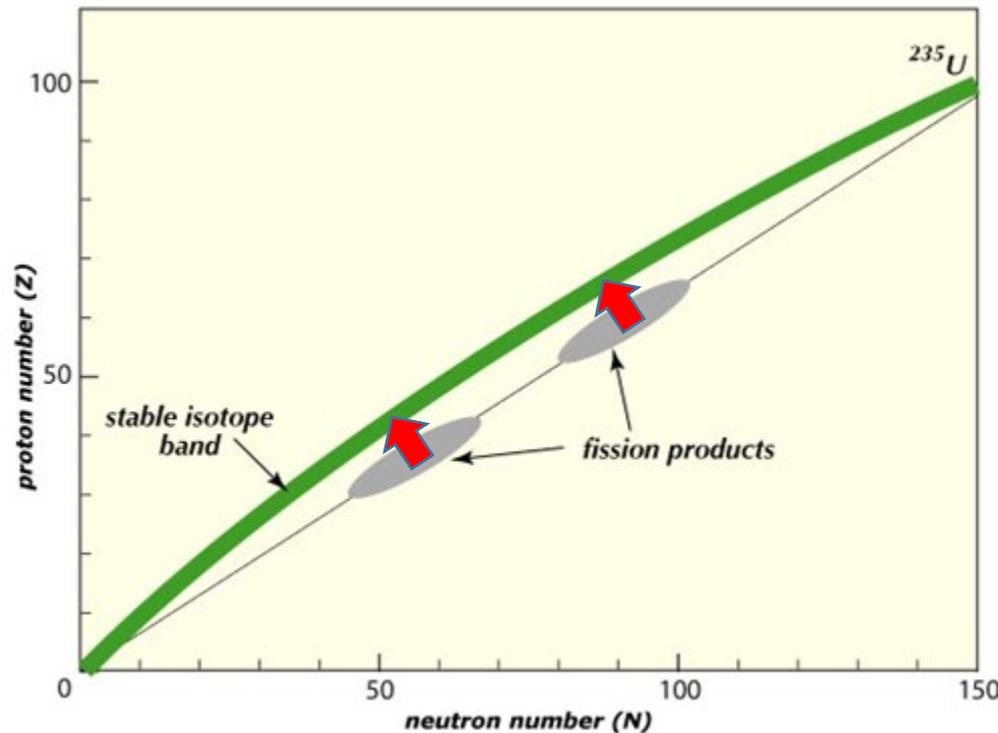
Alternative fission reactions

Nuclear Fuel Breeder



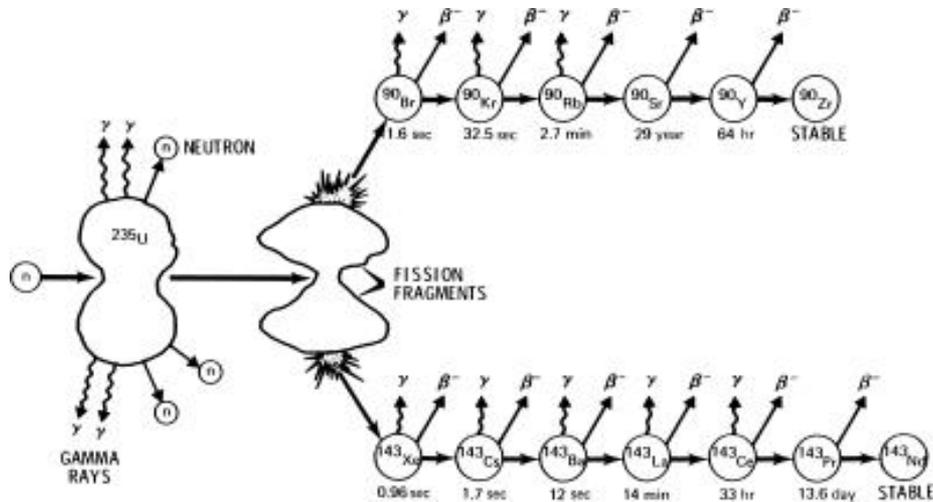
The breeding of nuclear fuel doesn't solve the issues associated with radioactive fission products but enhances the efficiency and reduced the fuel costs since ^{228}U and ^{232}Th is much more abundantly available than the rare ^{235}U . This removes an important costly aspect of fuel separation and preparation. Breeder also rely on fast neutrons and need no moderator materials.

Fission Process

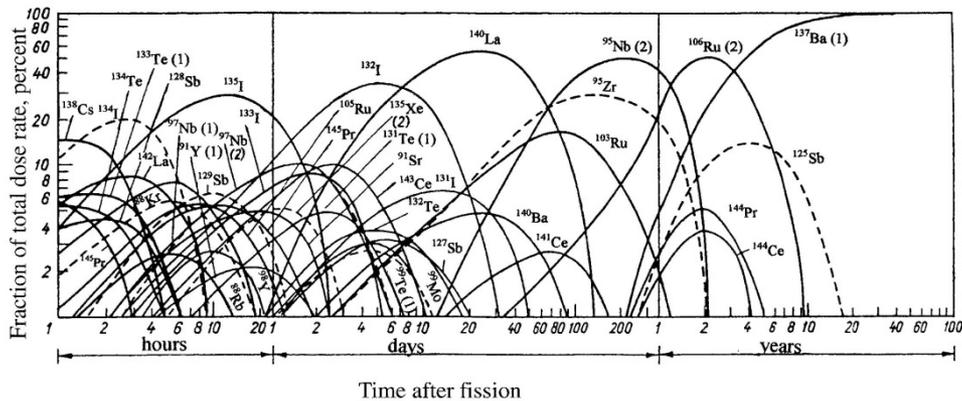


Fission product yields by mass for thermal neutron fission of ^{235}U , ^{239}Pu , a combination of the two typical of current nuclear power reactors, and ^{233}U used in the thorium cycle. The fission products are very neutron-rich and therefore highly radioactive and decay by β decay back to stability. Long-lived isotopes in the ^{233}U decay chains are: ^{93}Zr $T_{1/2}=1.5 \cdot 10^6$ y, ^{137}Cs $T_{1/2}=30$ y;

Fission products



Uranium fission and possible fission products in subsequent β^- decay processes along the isobaric lines towards stability. These decay chains originate from the primary fission product distribution.



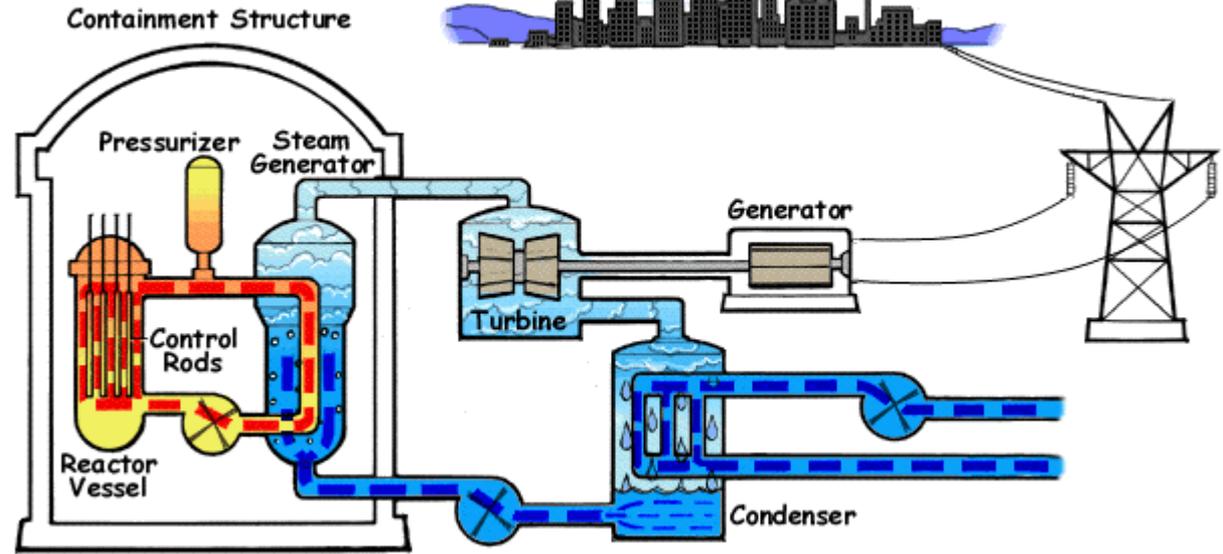
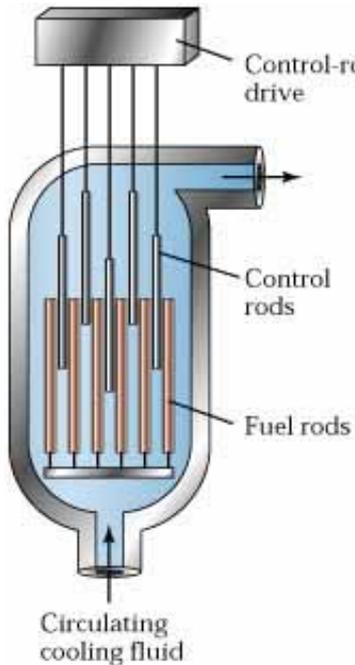
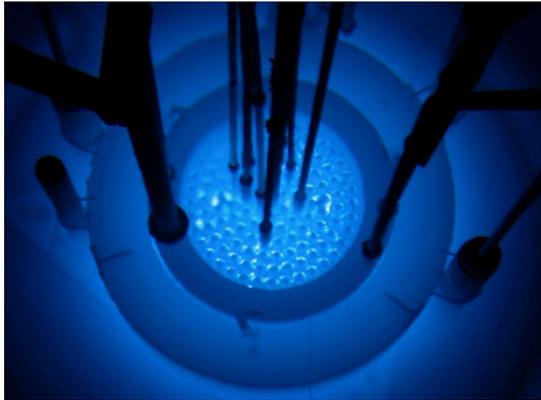
Medium-lived fission products

Prop: Unit:	$t_{1/2}$ (a)	Yield (%)	Heat (keV)	$\beta\gamma$
155Eu	4.76	0.0803	252	$\beta\gamma$
85Kr	10.76	0.2180	687	$\beta\gamma$
113mCd	14.1	0.0008	316	β
90Sr	28.9	4.505	2826	β
137Cs	30.23	6.337	1176	$\beta\gamma$
121mSn	43.9	0.00005	390	$\beta\gamma$
151Sm	96.6	0.5314	77	β

Long-lived fission products

Prop: Unit:	$t_{1/2}$ (Ma)	Yield (%)	Heat (keV)	$\beta\gamma$
99Tc	0.211	6.1385	294	β
126Sn	0.230	0.1084	4050	$\beta\gamma$
79Se	0.327	0.0447	151	β
93Zr	1.53	5.4575	91	$\beta\gamma$
135Cs	2.3	6.9110	269	β
107Pd	6.5	1.2499	33	β
129I	15.7	0.8410	194	β

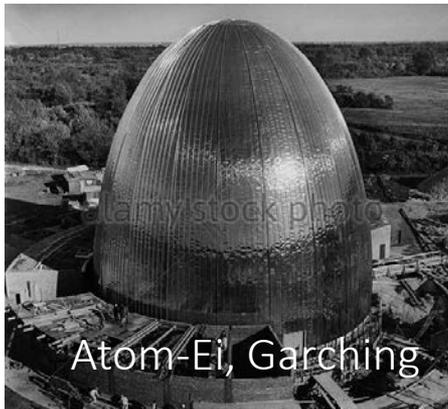
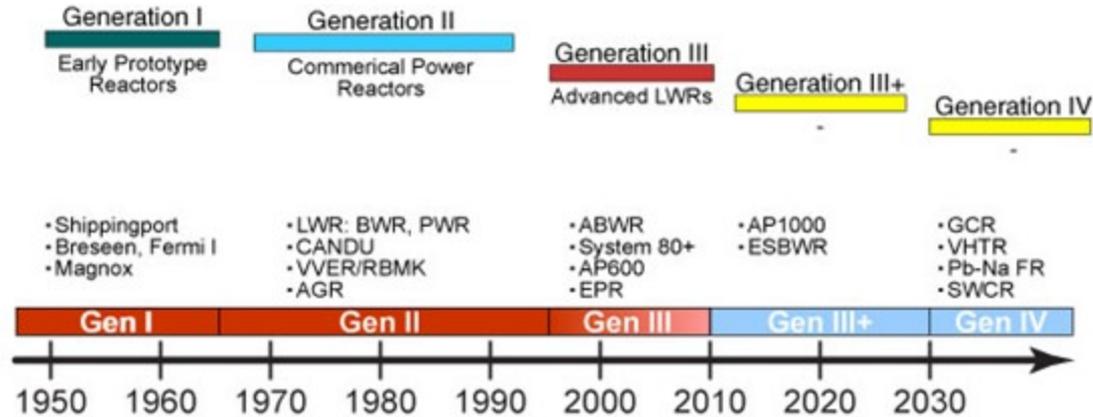
The basic technical principles



The cooling water that also functions as moderator for neutrons circulates through the reactor core and carries the heat that is generated by the fission process away in a closed water cycle, heating water in a secondary cycle. The water in the first cycle is radioactive from fission products the water in the second cycle should not be radioactive unless a leak has occurred. The hot water or steam in the second cycle drives the turbine. The control rods are made of Boron containing material since Boron has a large cross section for neutron capture. If in place the released fission neutrons are absorbed and the chain fission reactions stops. But there is still substantial decay heat being released, which requires continuous cooling!

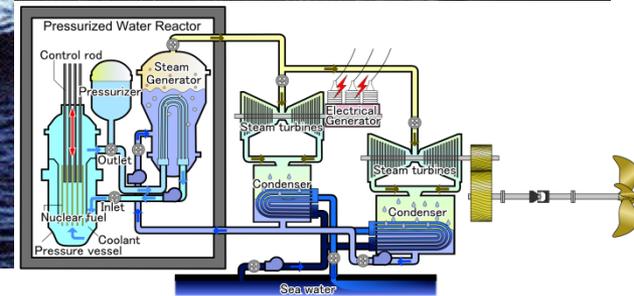
Reactor Generations

Technical developments and safety concerns – sometimes even new physics results - drive the change in the different reactor generations from the early prototypes such as the pile to latest generation of reactors with high neutron flux to high energy out-put. A reactor life-time is 40-50 years, Public and political concerns often lead to administrative delays forcing the lifetime extension with consequences in aging and safety reduction.



The pressurized light water reactor

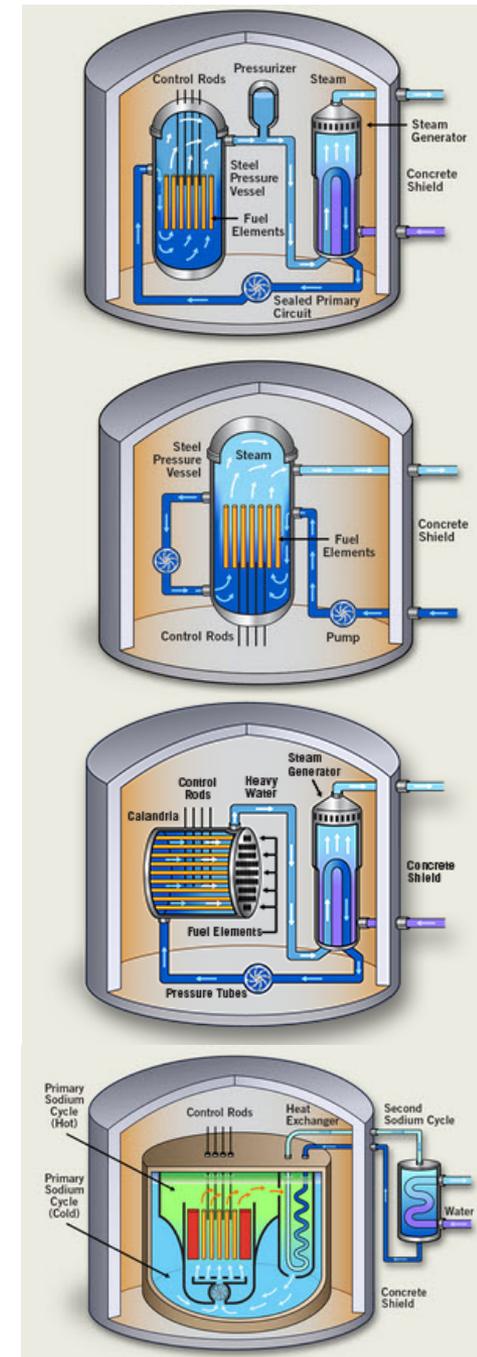
The PLWR is the present is the most used reactor type, developed at Oak Ridge and utilized for the nuclear submarine development under Admiral Rickover.



The Light Water Reactor (LWR) concept of water cooling seemed the obvious choice since water was available in large abundance. However, Rickover was also charged with the development of reactor types for peaceful applications (Atoms for Peace) and he continued with the same type despite the fact that it had more risk factors than other versions developed at Oak Ridge. The industrial military complex had started, since military requirements dictated the direction of civilian developments.

Reactor types and coolants

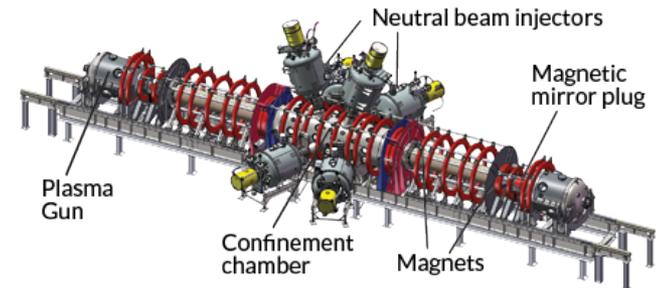
Reactor type	Fuel	Moderator	Coolant	Number
Pressurized water reactor (PWR)	Enriched UO_2	Water	Water	290
Boiling water reactor (BWR)	Enriched UO_2	Water	Water	78
Pressurized heavy water reactor (PHWR)	Natural UO_2	Heavy water	Heavy water	47
Light water graphite reactor (LWGR)	Enriched UO_2	Graphite	Water	15
Gas-cooled reactor (GCR)	Natural U, enriched UO_2	Graphite	Carbon dioxide	14
Fast breeder reactor (FBR)	PuO_2 and UO_2	None	Liquid sodium	3



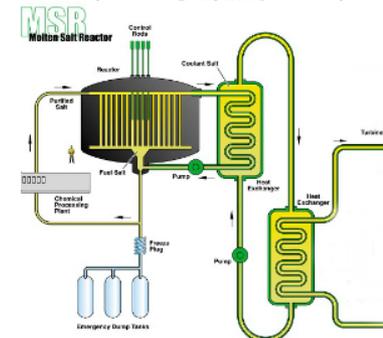
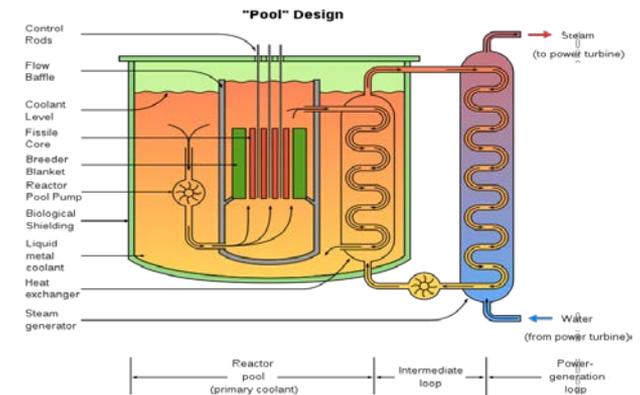
Other reactor types with potentially far advanced features

- **Boron Reactor**: a-neutronic fusion reactor device based on the $^{11}\text{B}(p,2\alpha)^4\text{He}$ reaction with little radioactive output from isotopic impurity $^{10}\text{B}(p,\alpha)^7\text{Be}(e^- \nu)^7\text{Li}$ ($T_{1/2}=53$ d). High intensity proton beam injected into a hot boron plasma in the magnetic confinement chamber
- **Liquid Metal Reactor (LMR)**: uses liquid sodium or lithium as coolant and uranium and/or thorium as fuel elements with a different distribution of fission products. Production of radioactive ^{24}Na occurs by neutron capture, but ^{24}Na has only 15 h half life. No stable ^8Li !
- **Molten Salt Reactor (MSR)**: the uranium fuel is dissolved in the sodium fluoride salt coolant which circulates through graphite core channels to achieve some moderation. Fission products are removed continuously and the actinides are fully recycled. A secondary coolant system is used for electricity generation.

Tri Alpha Energy model



Liquid Metal cooled Fast Breeder Reactors (LMFBR)



During Normal Operation

Material is being contained, in fuel rods and need to be separated in fuel facilities



Radioactivity accumulates in the fuel rods and may act as neutron poison – capture neutrons with subsequent γ -emission instead of fission. The efficiency of the fuel rod gradually declines and the fuel-rod has to be replaced and regenerated. This opens the problem of short-term storage and long-term deposition of radioactive waste.

Yield of Fission Products

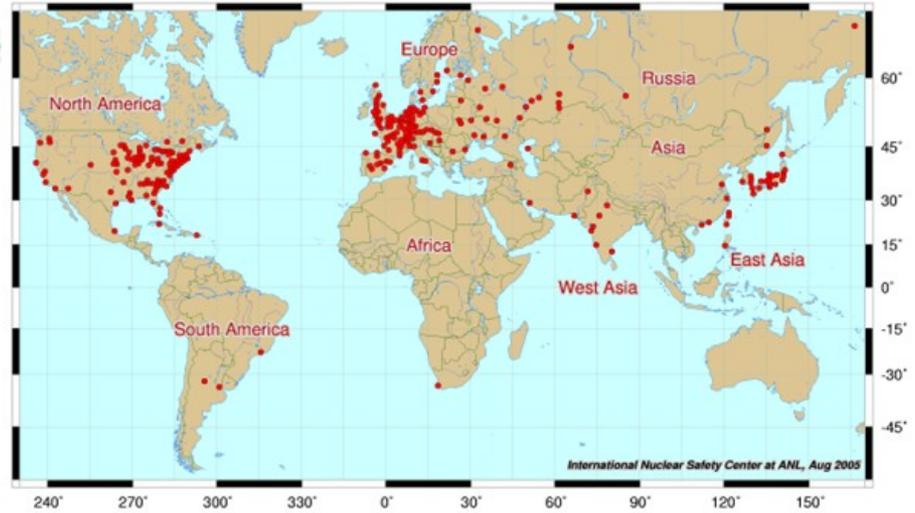
By yield

Yield	Element	Isotope	Half-life
6.7896%	Caesium	133Cs → 134Cs	2.065 y
6.3333%	Iodine, Xenon	135I → 135Xe	6.57 h
6.2956%	Zirconium	93Zr	1.53 My
6.1%	Molybdenum	99Mo	65.94 h
6.0899%	Caesium	137Cs	30.17 y
6.0507%	Technetium	99Tc	211 ky
5.7518%	Strontium	90Sr	28.9 y
2.8336%	Iodine	131I	8.02 d
2.2713%	Promethium	147Pm	2.62 y
1.0888%	Samarium	149Sm	virtually stable
0.9% ^[2]	Iodine	129I	15.7 My
0.4203%	Samarium	151Sm	90 y
0.3912%	Ruthenium	106Ru	373.6 d
0.2717%	Krypton	85Kr	10.78 y
0.1629%	Palladium	107Pd	6.5 My
0.0508%	Selenium	79Se	327 ky
0.0330%	Europium, Gadolinium	155Eu → 155Gd	4.76 y
0.0297%	Antimony	125Sb	2.76 y
0.0236%	Tin	126Sn	230 ky
0.0065%	Gadolinium	157Gd	stable
0.0003%	Cadmium	113mCd	14.1 y

By half-life

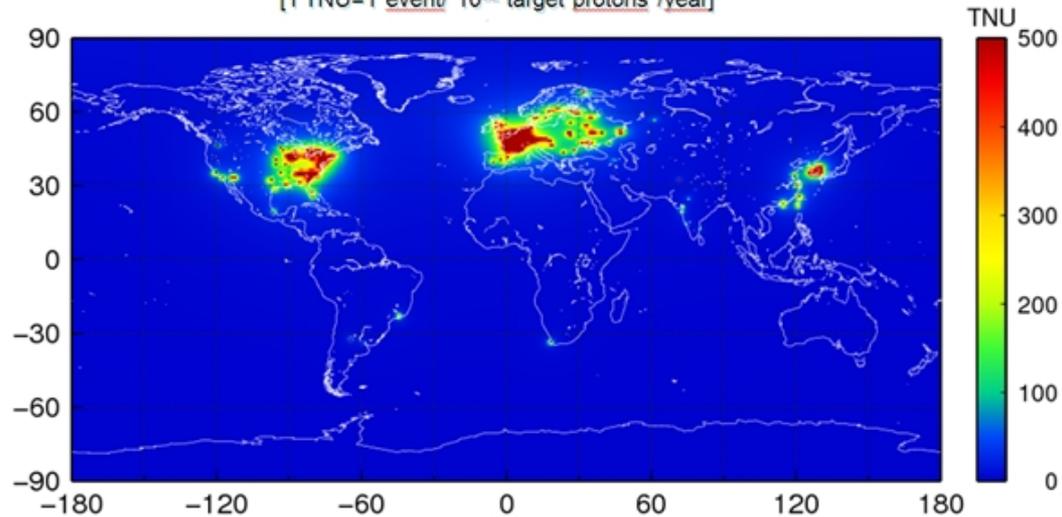
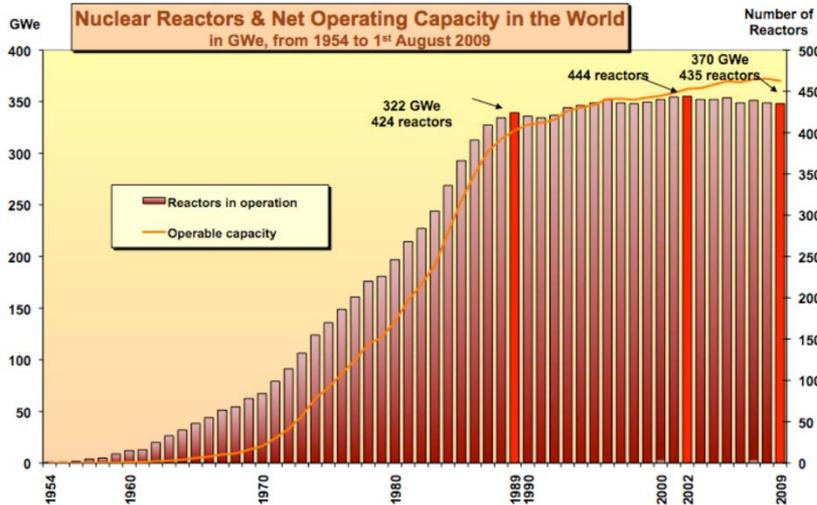
Yield	Isotope	Half-life
2.8336%	131I	8.02d
0.3912%	106Ru	373.6d
6.7896%	133Cs → 134Cs	2.065y
2.2713%	147Pm	2.62y
0.0297%	125Sb	2.76y
<0.0330%	155Eu → 155Gd	4.76y
0.2717%	85Kr	10.78y
<0.0003%	113mCd	14.1y
5.7518%	90Sr	28.9y
6.0899%	137Cs	30.17y
<0.4203%	151Sm	90y
6.0507%	99Tc	211ky
0.0236%	126Sn	230ky
0.0508%	79Se	327ky
6.2956%	93Zr	1.53My
<6.3333%	135Cs	2.3My
0.1629%	107Pd	6.5My
0.6576%	129I	15.7My
<1.0888%	149Sm	nonradioactive
<0.0065%	157Gd	nonradioactive

Reactor Locations, US and World-Wide

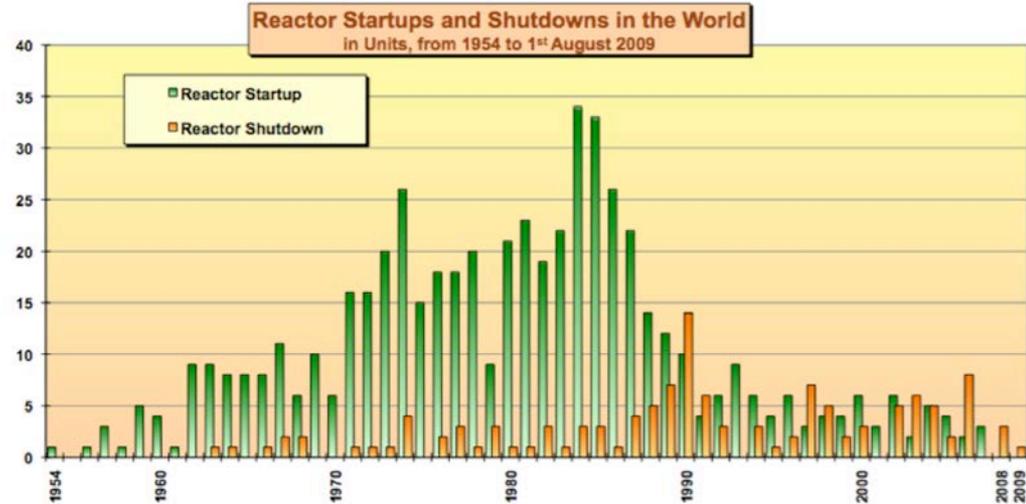
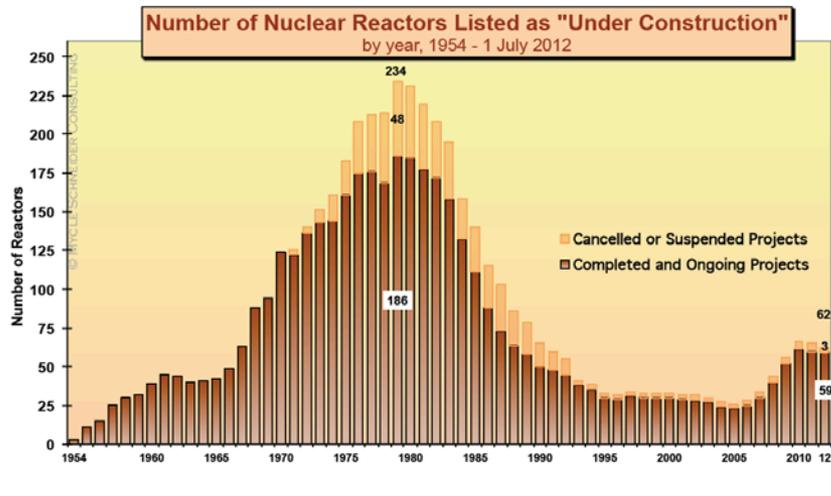


Map of neutrino flux

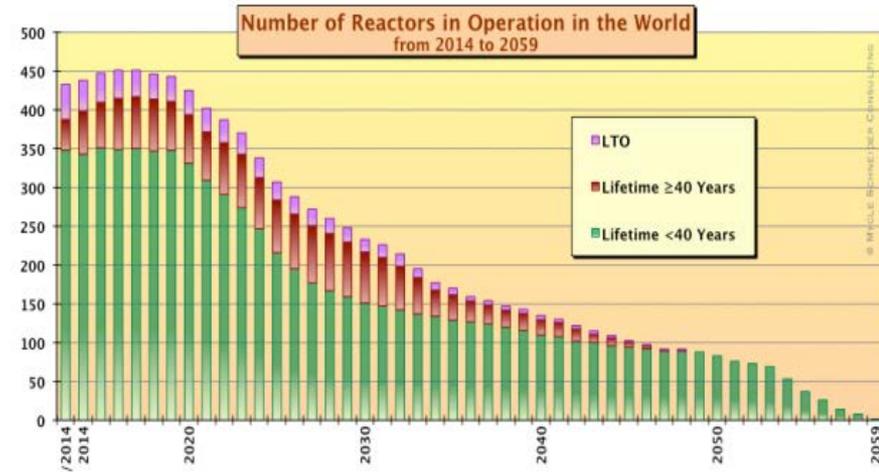
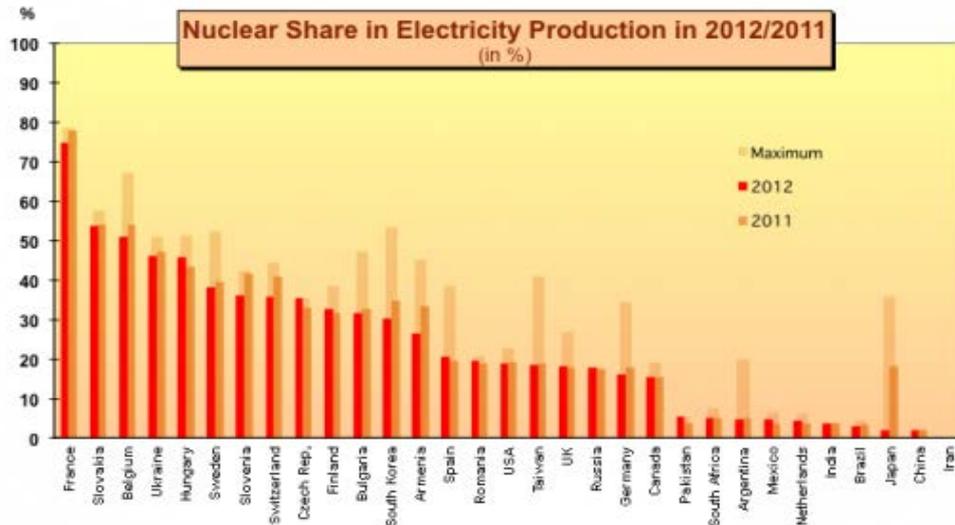
[1 TNU=1 event/ 10^{32} target protons /year]



Nuclear Reactor Budgeting



Clear indication of decline in nuclear energy production, peak in 1975-1980



Nuclear Plant Decommissioning



“DECOMMISSIONING”
a nice word for a super expensive
and almost impossible job



Radioactive for 100,000 years