## Nuclear Reactions Experiment - I



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## OVERVIEW

Lecture 1 Focus: Overview of Nuclear Reactions
Lecture 2 Focus: Example using single-particle transfer reactions

- Understanding different types / forms of reactions is key
- Integrated into measurements, New physics directly, Isotope production
- Reaction Formalism
- Overview of some reaction types
- Example cases of complementary reactions


## Lecture 1 Takeaways:

- Familiarity with reaction "language"
- List various reaction types \& their general properties
- Link reaction method and/or probe to physics quantities of interest
- How multiple reactions are used to explore common physics goals


## REACTIONS LINK VARIOUS ASPECTS OF NUCLEAR SCIENCE



## REACTIONS LINK VARIOUS ASPECTS OF NUCLEAR SCIENCE



Methods / Tools

## REACTIONS LINK VARIOUS ASPECTS OF NUCLEAR SCIENCE



## REACTIONS LINK VARIOUS ASPECTS OF NUCLEAR SCIENCE



## ISOTOPE DISCOVERY

## Leveraging various reaction types / energies / facilities

THE NUCLIDE TRAIL
Isotope discovery over the past 100 years (below) has jumped with each introduction of new technology. Some 2,700 radioactive isotopes have been discovered so far (below right), but about 3,000 more are predicted to exist.

## Isotope-discovery technique

Light particle reactionsNeutron reactions FusionFragmentation/spallation



## REACTION TYPES / ENERGIES / PHYSICS MOTIVATION <br> Many methods with overlapping science goals



## BREAKDOWN IN COULOMB BARRIER REGION Physics goals may be encompassed by various reaction types



## SOME BASIC REACTION FORMALISM / NOTATION

## Pragmatic way of writing nuclear reactions

The ingredients

- Target (A)
- Projectile (a)
- Beam-like outgoing ion (b)
- Target-like outgoing ion (recoil) (B)

Other Considerations


- Inverse kinematics [rare-isotope beams, $\mathrm{MeV} / \mathrm{u}$ ]
- near-Coulomb energies
- Low energy (<20 MeV/u) to Intermediate energy (50-few
 hundred $\mathrm{MeV} / \mathrm{u}$ )

For most reactions it is the $(a, b)$ of $A(a, b) B$ that is used to label the reaction

## REACTION TYPES

## TYPES OF NUCLEAR REACTIONS

## Beam energies range from $\mathrm{eV} / \mathrm{u}$ to $\mathrm{GeV} / \mathrm{u}$

- Direct reactions
- Knockout - single / multi-particle
- Transfer - single / multi-nucleon / charge-exchange
- Scattering
- Inelastic / Elastic / Resonance
- Fusion
- Compound - Evaporation / Fission
- Capture
- Neutron / Proton / Alpha / Induced reactions
- Others
- Heavy-ion collisions / Fragmentation / Deep inelastic collisions / etc...


## FUSION REACTIONS

## Fusion \& fusion-fission: ${ }^{12} \mathrm{C}\left({ }^{( } \mathrm{C}, \mathrm{x}\right), \mathrm{Th}\left({ }^{15} \mathrm{C}, \mathrm{X}\right)$




## KEY POINTS

- <10 MeV/u
- No "memory" of formation
- Size \& shape of barrier
- Relevance to stellar processes \& heavyelement creation


Extract: excitation functions, angular momenta
Deduce: barrier heights, fusion probabilities, S-factors
Tools: various cross section codes [HF approach / PACE / CASCADE / HIVAP / etc...]

## FUSION REACTIONS

## Fusion evaporation: ${ }^{20} \mathrm{Ne}\left({ }^{22} \mathrm{Ne}, \mathrm{a} 2 \mathrm{n}\right){ }^{36} \mathrm{Ar}$

## KEY POINTS

- Compound has no "memory" of construction
- Population along yrast line
- Provides alignment
- Prolific tool in gamma-ray spectroscopy

Extract: Level \& decay schemes, angular momenta, transition strengths
Deduce: nuclear shapes, entry distributions, deformation parameters
Tools: various fusion-evap codes [PACE / CASCADE / HIVAP / etc...]

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## SCATTERING REACTIONS <br> Coulomb excitation \& Coulomb dissociation

## KEY POINTS

- Safe < Coulomb barrier $\sim 3 \mathrm{MeV} / \mathrm{u}$
- Intermediate > Coulomb barrier
- Dissociation > 250 MeV/u
- Electromagnet probe (virtual photon flux)
- Deformation
- Capture rates
- Equipped for inverse kinematics
$\sigma_{\pi \lambda} \approx\left(\frac{Z_{\text {pro }} e^{2}}{\hbar c}\right)^{2} \frac{\pi}{e^{2} b_{\text {min }}^{2 \lambda-2}} B(\pi \lambda, 0 \rightarrow \lambda) \begin{cases}1 /(\lambda-1) & \text { for } \lambda \geqslant 2 \\ 2 \ln \left(b_{a} / b_{\text {min }}\right) & \text { for } \lambda=1\end{cases}$


## Excitation of ${ }^{34} \mathrm{Si}$ to ${ }^{34}{ }^{3}{ }^{*}$ :

Extract: transition strengths
Deduce: deformation parameters

Reaction Rates for ${ }^{26} \mathrm{Si}(\mathrm{p}, \mathrm{p})$ from dissociation ${ }^{27} \mathrm{P}$ to ${ }^{26} \mathrm{Si}+\mathrm{p}$ :
Extract: Survival probabilities, El strengths Deduce: capture rates

## SCATTERING REACTIONS <br> Coulomb excitation \& Coulomb dissociation

## KEY POINTS

- Safe < Coulomb barrier ~3 MeV/u

$$
N_{A}\langle\sigma v\rangle^{\mathrm{tot}}=\sum_{i} N_{A}\langle\sigma v\rangle_{i}^{\mathrm{res}}+N_{A}\langle\sigma v\rangle^{\mathrm{dc}}
$$


$\sigma_{\pi \lambda} \approx\left(\frac{Z_{\text {pro }} e^{2}}{\hbar c}\right)^{2} \frac{\pi}{e^{2} b_{\min }^{2 \lambda-2}} B(\pi \lambda, 0 \rightarrow \lambda) \begin{cases}1 /(\lambda-1) & \text { for } \lambda \geqslant 2 \\ 2 \ln \left(b_{a} / b_{\min }\right) & \text { for } \lambda=1\end{cases}$

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## SCATTERING REACTIONS <br> Elastic \& resonance scattering (p,p), (d,d), ( $\alpha, \alpha$ )



## KEY POINTS

- Few to hundreds $\mathrm{MeV} / \mathrm{u}$
- Sensitive to probe: proton, deuteron, alpha, etc.
- Scan in angle and/or energy
- Resonance params on analog states
- sp structure
- Widths
- Key in development of optical model description of nuclear potential

Extract: angular distributions, analog states
Deduce: sp states, optical model parameters, resonance/decay widths
Tools: R-Matrix, optical model calculations [DWBA, coupled channels, etc...]

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## SCATTERING REACTIONS

Inelastic scattering: ( $\left.p, \mathrm{p}^{\prime}\right),\left(\mathrm{d}, \mathrm{d}^{\prime}\right),\left(\mathrm{a}, \mathrm{a}^{\prime}\right),\left({ }^{12} \mathrm{C},{ }^{12} \mathrm{C}^{\prime}\right)$

## KEY POINTS

- Few to 100 's of MeV/ u
- Selective to probe:
- Proton - isovector
- Deuteron isoscalar
- Resonance structures
- Cluster structures
- Collective features in nuclei

Extract: Distributions, resonance strengths
Deduce: unique excitation modes, clustering prob., isoscalar / isovector modes, deformation length Tools: R-Matrix, optical model calculations [DWBA, CC, etc...]

## SCATTERING REACTIONS

Inelastic scattering: (p,p), (d, d'), ( $\left.\alpha, \alpha^{\prime}\right),\left({ }^{12} \mathbf{C}^{12} \mathbf{}^{\prime}{ }^{\prime}\right)$

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## HEAVY-ION COLLISIONS <br> ${ }^{A} \mathrm{Sn}+{ }^{\text {ASn }}, \mathrm{ANi}+{ }^{\text {ANi }}$



## KEY POINTS

- >350 MeV/u
- Pion production threshold >280 MeV
- Hot dense matter
- Stars
- Pion production as a test of the symmetry energy


Extract: pion production, +/- asymmetry Deduce: Symmetry energy, equation of state


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## DIRECT REACTIONS

Single-nucleon transfer: (d,p), ( ${ }^{(3 \mathrm{He}, \mathrm{d}),\left({ }^{13} \mathrm{C},{ }^{12} \mathrm{C}\right)}$

## KEY POINTS

- ~3-20 MeV/u
- Highly selective
- Direct probe of single-particle aspects
- Surrogate ( $p, \gamma$ ) / ( $n, \gamma$ )
- resurgence in the RIB era
- Beam production


Extract: orbital angular momenta, spectroscopic overlaps, energy centroids Deduce: nucleon occupancies, single-particle energies, two-body matrix elements Tools: Distorted wave Born approximation [DWBA], Coupled Channels, etc...]

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## DIRECT REACTIONS

## Multi-nucleon transfer e.g., (p,t), (7Li,t), (6Li,p)



## KEY POINTS

- ~3-20 MeV/u
- selective
- Alpha-like transfer: ( $\alpha, \mathrm{y}$ ), ( $\alpha, X$ )
- Sensitive to paring (2n)
- Exploratory - cluster / rotational states
- resurgence in the RIB era

Extract: final state angular momenta, spectroscopic overlaps, resonance widths Deduce: resonance strengths, reaction rates, pair occupancies, collectivity
Tools: Distorted wave Born approximation [DWBA], Coupled Channels, etc...]

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$\frac{0^{+}, 69 \%}{{ }^{96} \mathrm{Mo}_{54}} \frac{0^{+}, 86 \%}{{ }^{98} \mathrm{Mo}_{56}} \quad \frac{0^{+}, 100 \%}{{ }^{100} \mathrm{Mo}_{58}}{ }^{102} \mathrm{Mo}_{60}{ }^{0^{+}, 77 \%} \quad \frac{0^{+}}{{ }^{104} \mathrm{Mo}_{62}} \frac{0^{+}}{{ }^{106} \mathrm{Mo}_{64}}$
Transitions strengths normalised to ${ }^{100} \mathrm{Mogs}$.


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## DIRECT REACTIONS

Charge-exchange: (p,n), (7Li, $\left.{ }^{7} \mathrm{Be}\right)$

## KEY POINTS

- ~5-400 MeV/u
- Isobaric analog states

- Gamow-Teller distributions
- Astrophysics
- Neutrino physics
- Beam production method


Extract: angular distributions, isobaric analog states
Deduce: Gamow-Teller strength distributions, level densities, $g$-strength functions
Tools: Distorted wave Born approximation [DWBA], Coupled Channels, etc...]

## DIRECT REACTIONS

Charge-exchange: (p,n), (7Li, ${ }^{7 B e}$ )


$$
\mathrm{B}\left(\mathrm{GT}_{ \pm}\right)=\frac{1}{2 J_{i}+1}\left|\left\langle\Psi_{f}\left\|\sum_{j=1}^{A} \sigma_{j} \boldsymbol{\tau}_{ \pm, j}\right\| \Psi_{i}\right\rangle\right|^{2}
$$

## KEY POINTS

- ~5-400 MeV/u
- Isobaric analog states
- Gamow-Teller distributions
- Astrophysics
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Initial nucleus


Target

Initial nucleus


Final nucleus


Final nucleus

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Extract: angular distributions, isobaric analog states


Final nucleus



Excitation Energy [MeV]

Deduce: Gamow-Teller strength distributions, level densities, $g$-strength functions
Tools: Distorted wave Born approximation [DWBA], Coupled Channels, etc...]

## DIRECT REACTIONS <br> Quasi-free \& nucleon knockout: (p,2p), (9Be, -2p)

## KEY POINTS

- >50 MeV/u knockout
- >350 MeV/u quasi-free knockout
- selective to hole states

- Study of overlaps w/ established tools
- Pairing force
- Efficient in the RIB era


Extract: orbital angular momenta, spectroscopic overlaps
Deduce: occupancies, single-particle energies, pairing strengths
Tools: Eikonal \& Glauber model, impulse approximation, Coupled channels calculations, etc...

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residue moment distribution
$\rightarrow \ell$-value of knocked-out $n$


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Initial nucleus [beaml


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## INTERMEDIATE ENERGY FRAGMENTATION

Create \& populate isotopes at the extremes >100 MeV/u

## DEEP INELASTIC REACTIONS

Production of exotic nuclei via multi-nucleon removal + exchange


## CAPTURE REACTIONS

p,n,o cross sections key to proliferation \& astrophysics


## COMPLEMENTARITY OF NUCLEAR REACTIONS



## ISOMERIC BEAM PRODUCTION

Desire a beam of ${ }^{34} \mathrm{Cl}$ residing in either (or both) isomeric \& ground states
$34 \mathrm{~g}, \mathrm{mCl}(\mathrm{p}, \mathrm{\gamma})$ rates influence ${ }^{34} \mathrm{~S}$ production in classical novae impacting solar grain classification

A=35 Mirror Pair
Partial Level \& Decay Schemes

$34 \mathrm{~g}, \mathrm{~m} \mathrm{Cl}(\mathrm{d}, \mathrm{p})$ mirror reaction is of interest

ISOMERIC BEAM PRODUCTION
Sub-set of beam production options
$34 m, g \mathrm{Cl}(\mathrm{Z}=17, \mathrm{~N}=17)$


## ISOMERIC BEAM PRODUCTION

Sub-set of beam production options

## Transfer Reactions



- proton adding: (d,n), (3He,d)
- Neutron removal: (d,t), (3He, a )
- Charge exchange: (p,n)
- Multi-nucleon: (d, $\alpha$ ), ( $\alpha, d$ ), ( $\alpha, n$ )


ISOMERIC BEAM PRODUCTION

Sub-set of beam production options


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- Neutron removal: (d,t), (3He, a )
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## ISOMERIC BEAM PRODUCTION

Sub-set of beam production options
Fragmentation $>100 \mathrm{MeV} / \mathrm{u}$


- ${ }^{36} \mathrm{Ar}+\mathrm{Be}$ : pn removal
- ${ }^{40} \mathrm{Ca}+\mathrm{Be}$ apn removal



## ISOMERIC BEAM PRODUCTION

## Sub-set of beam production options

## Fusion Evaporation



- ${ }^{16} \mathrm{O}(20 \mathrm{Ne}, \mathrm{pn}),{ }^{24} \mathrm{Mg}\left({ }^{(12} \mathrm{C}, \mathrm{pn}\right),{ }^{27} \mathrm{Al}\left({ }^{12} \mathrm{C}, \mathrm{an}\right)$



## ISOMERIC BEAM PRODUCTION

## Sub-set of beam production options

## Fusion Evaporation



- ${ }^{16} \mathrm{O}(20 \mathrm{Ne}, \mathrm{pn}),{ }^{24} \mathrm{Mg}\left({ }^{(12} \mathrm{C}, \mathrm{pn}\right),{ }^{27} \mathrm{Al}\left({ }^{12} \mathrm{C}, \mathrm{an}\right)$

Other Possible Reactions:

- Resonance scattering ${ }^{33} S(p, \gamma)$
- Electron capture
- Spallation


## KEY ASTROPHYSICS REACTIONS RATES

Requires knowledge of resonance energies, spins, widths (overlaps), ...
thermonuclear reaction rate:
$\langle\sigma v\rangle=\sqrt{\frac{8}{\pi \mu}} \frac{1}{(k T)^{3 / 2}} \int_{0}^{\infty} \sigma(E) E e^{-E / k T} d E$ :
partial width

$$
\Gamma_{x}=C^{2} S_{x} \times \Gamma_{x}^{s . p .}
$$

s-factor


C

$$
S(E)=\sigma(E) \times E \times e^{2 \pi \eta}
$$

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resonance capture

$$
\sigma(E)=\frac{\lambda^{2}}{4 \pi} \frac{2 J_{C^{*}}+1}{\left(2 J_{A}+1\right)\left(2 J_{x}+1\right)} \frac{\Gamma_{x} \Gamma_{y}}{\left(E-E_{r}\right)^{2}+\Gamma^{2} / 4}
$$

partial width

$$
\Gamma_{x}=C^{2} S_{x} \times \Gamma_{x}^{s . p .}
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s-factor


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S(E)=\sigma(E) \times E \times e^{2 \pi \eta}
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## 14O(A,P) ${ }^{17 F}$ REACTION RATE hot-CNO breakout in type-I x-ray bursts




EXAMINATION OF THE ROLE OF THE ${ }^{14} \mathrm{O}(\alpha$,


PHYSICAL REVIEW C 90,025803 (2014)

| $E_{x}(\mathrm{MeV})^{n}$ | $E_{\text {res }}(\mathrm{MeV})^{\text {n }}$ | $J^{*}$ | $\Gamma_{\alpha}(\mathrm{eV})$ | $\Gamma_{P}(\mathrm{keV})$ | $\Gamma_{p^{\prime}}(\mathrm{keV})$ | $\Gamma(\mathrm{keV})$ | $\omega \gamma$ ( MeV ) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $5.153 \pm 0.01$ | 0.039 | $3{ }^{-}$ | $4.3 \times 10^{-52 a}$ | $1.7{ }^{\text {a }}$ |  | $\leqslant 15^{\text {a }}$ | $3.0 \times 10^{-57}$ |
| $6.150 \pm 0.01$ | 1.036 | $1^{-}$ | $3.9 \pm 1.0^{\text {b }}$ | $37.8 \pm 1.9{ }^{\text {c }}$ | $15.9 \pm 0.7^{c}$ | $53.7 \pm 2.0^{\text {c }}$ | $1.2 \times 10^{-5}$ |
| $6.286 \pm 0.01$ | 1.172 | $3-$ | $0.34{ }^{\text {a }}$ | $20 \pm 15^{\text {d }}$ |  | $20 \pm 15^{\text {d }}$ | $2.4 \times 10^{-6}$ |
| $7.05 \pm 0.03$ | 1.936 | $4^{+}$ | $22.6 \pm 3.2^{\text {e }}$ | $90 \pm 40^{5}$ |  | $90 \pm 40^{f}$ | $2.0 \times 10^{-4}$ |
| $7.35 \pm 0.02$ | 2.236 | $2^{+}$ | $40 \pm 30^{\prime}$ | $70 \pm 60^{f}$ |  | $70 \pm 60$ f | $2.0 \times 10^{-4}$ |
| $7.62 \pm 0.02$ | 2.506 | $1-$ | $1000 \pm 120^{\circ}$ | $72 \pm 20^{t}$ | $<2^{\text {f }}$ | $75 \pm 20$ ? | $3.0 \times 10^{-3}$ |
| $7.94 \pm 0.01$ | 2.826 | $3^{-}$ | $(11 \pm 6.6) \times 10^{7 \mathrm{E}}$ | $35 \pm 15^{2}$ | $9.0 \pm 5.6^{2}$ | $55 \pm 20^{8}$ | $6.2 \times 10^{-2}$ |

${ }^{14} \mathrm{O}(\mathrm{A}, \mathrm{P}){ }^{17} \mathrm{~F}$ REACTION RATE hot-CNO breakout in type-I x-ray bursts

## Exploring the mirror states in ${ }^{18} \mathrm{O}$

- ${ }^{17} \mathrm{O}(\mathrm{d}, \mathrm{p})$ : neutron transfer
- C²S values of mirror state, $E, \pi$
- ${ }^{16} \mathrm{O}(\mathrm{t}, \mathrm{p})$ : 2 n transfer
- E, $\pi$ of mirror levels
- ${ }^{14} \mathrm{C}\left({ }^{6} \mathrm{Li}, \mathrm{d}\right)$ : alpha transfer
- Alpha width of mirror state

${ }^{14} \mathrm{O}(\mathrm{A}, \mathrm{P}){ }^{17} \mathrm{~F}$ REACTION RATE hot-CNO breakout in type-I x-ray bursts



## Multi-particle transfer reactions

- ${ }^{16} \mathrm{O}(3 \mathrm{He}, \mathrm{n}): 2 \mathrm{p}$ transfer
- ${ }^{12} \mathrm{C}\left({ }^{12} \mathrm{C}, 6 \mathrm{He}\right)$ : exotic transfer
- ${ }^{20} \mathrm{Ne}(\mathrm{p}, \mathrm{t}):-2 \mathrm{n}$ removal
- Selective reactions
- Determine resonance E, total widths (Г), L(J), $\pi$



## ${ }^{14} \mathrm{O}(\mathrm{A}, \mathrm{P}){ }^{17} \mathrm{~F}$ REACTION RATE hot-CNO breakout in type-I x-ray bursts



Elastic / resonance scattering

- 17F(p,p): inverse proton scattering
- (p,p’) contributions
- Determine resonance E, partial widths (Г), L(J), п
- R-Matrix interpretation

${ }^{14} \mathrm{O}(\mathrm{A}, \mathrm{P}){ }^{17} \mathrm{~F}$ REACTION RATE hot-CNO breakout in type-I x-ray bursts



## Elastic / resonance scattering

- ${ }^{17} \mathrm{~F}(\mathrm{p}, \mathrm{p})$ : inverse proton scattering
- (p,p’) contributions
- Determine resonance E, partial widths (Г), L(J), п
- R-Matrix interpretation



## SINGLE-PARTICLE STRENGTHS <br> Complementary reactions: ${ }^{136} \mathrm{Xe}(\mathrm{p}, \mathrm{p})$ and ${ }^{136} \mathrm{Xe}(\mathrm{d}, \mathrm{p})$



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## CONCLUDING REMAKES ON LECTURE I

- Understanding different types / forms of reactions is key
- Integrated into measurements, New physics directly, Isotope production
- Reaction Formalism
- Overview of some reaction types
- Example cases of complementary reactions


## Lecture 1 Takeaways:

- Familiarity with reaction "language"
- List various reaction types \& their general properties
- Link reaction method and/or probe to physics quantities of interest
- How multiple reactions are used to explore common physics goals


## RESOURCES <br> Less than complete set of links / references: crhoffman@anl.gov

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Department of Physics, School of Physics and Chemistry, University of Surrey, Guildford, Surrey, GU2 7XH, United Kingdom

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C.R. Brune. ${ }^{1}$ W.H. Geist. ${ }^{1, *}$ R. W. Kavanagh, ${ }^{2}$ and K.D. Veal ${ }^{1, *}$

University of North Caroliza, Chapel Hill, North Carolina 27599-3255
Triangle Universities Nuclear Laboratory, Durham, North Carolina 27708-0308 'W.K. Kellogg Radiation Laboratory, Calljorouna Inststutere of fechnology, Pasadena, California 91125
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 Paul, and C Ugalde
Phys. Rev. Lett 112. 192701 - Publisheo 14 May 2014

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## LETTER

An increase in the ${ }^{12} \mathrm{C}+{ }^{12} \mathrm{C}$ fusion rate from resonances at astrophysical energies


## RESOURCES <br> Less than complete set of links / references: crhoffman@anl.gov

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## THE END

