

A Brief History

of

EBs @ ND

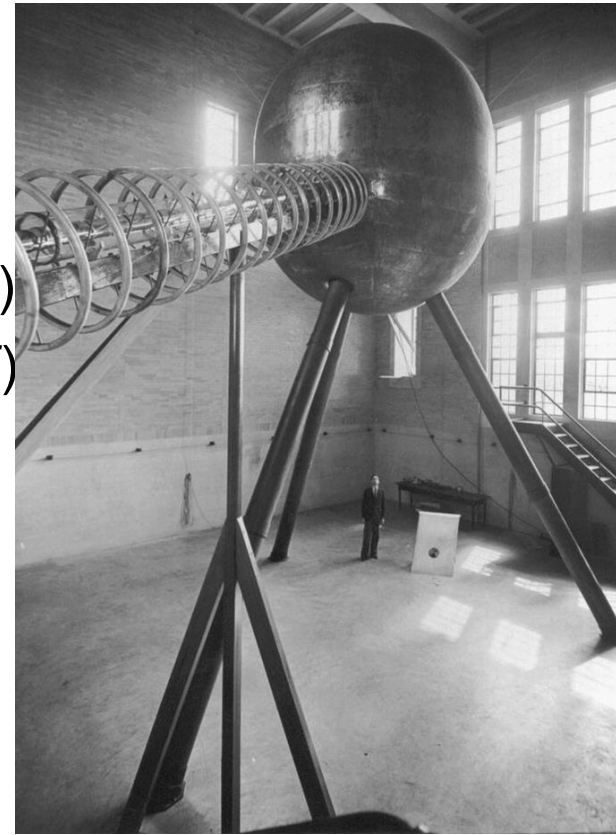
from

1936 – 2022

J.J. Kolata  
Exotic Beam Summer School  
University of Notre Dame  
June 7, 2022

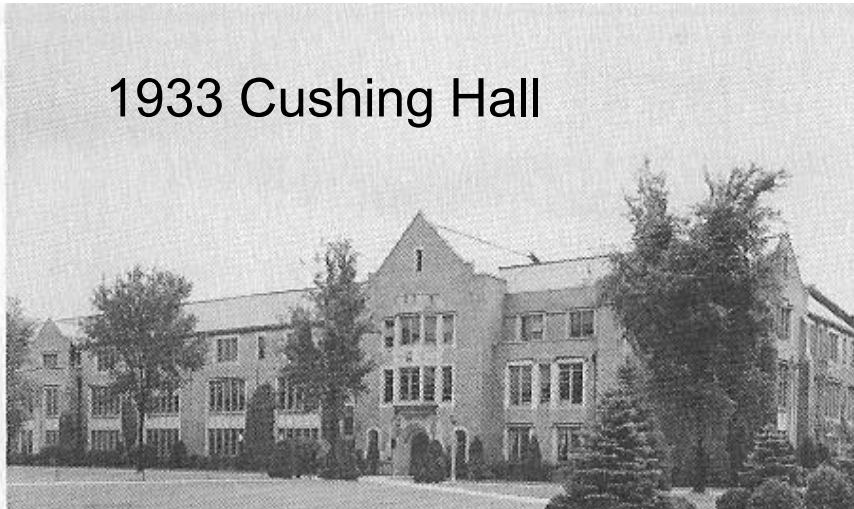
# Accelerator History

- 1930 Cockroft-Walton (Cavendish Laboratory)
- 1931 Converted X-ray tube (Caltech) (C.C. Lauritsen)
- 1931 Electrostatic Generator (MIT) (R. Van de Graaff)
- 1932 Cyclotron (UC Berkeley) (T.H. Lawrence)
- 1933 Electrostatic Generator (Carnegie) (M.A. Tuve)
- 1935 Electrostatic Generator (Carnegie) (R.G. Herb)
- 1936 Electrostatic Generator (Notre Dame)



The 6<sup>th</sup> accelerator laboratory in the world!

1933 Cushing Hall



# First RNB Apparatus (1988)

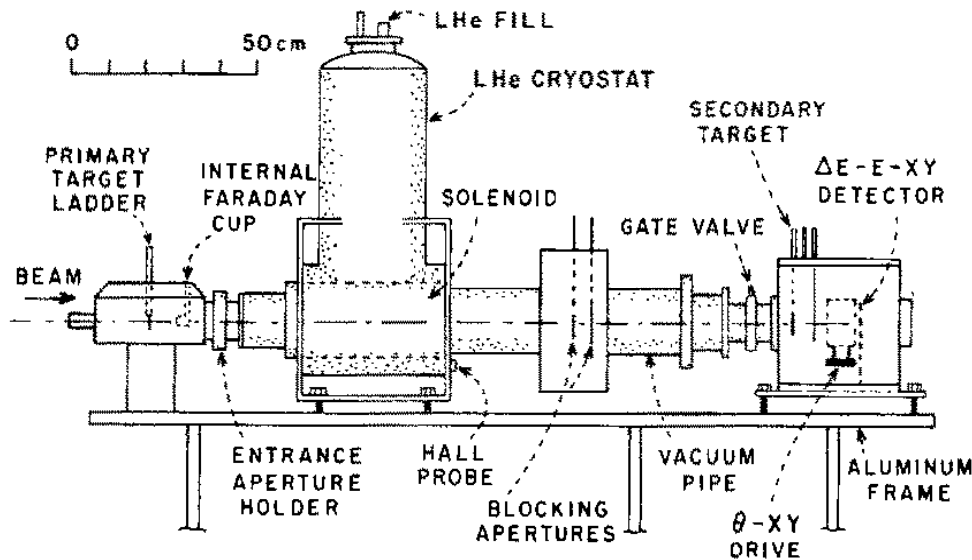


Fig. 1. Experimental apparatus for radioactive beam production.

Solenoid: 3.5 T ; 20 cm bore ; 40 cm length

Beams: "3-stage" VdG at Notre Dame

Nucl. Inst. Methods B40/41, 503 (1989)

Collaboration: J.J. Kolata (Notre Dame)  
F.D. Becchetti (Michigan)

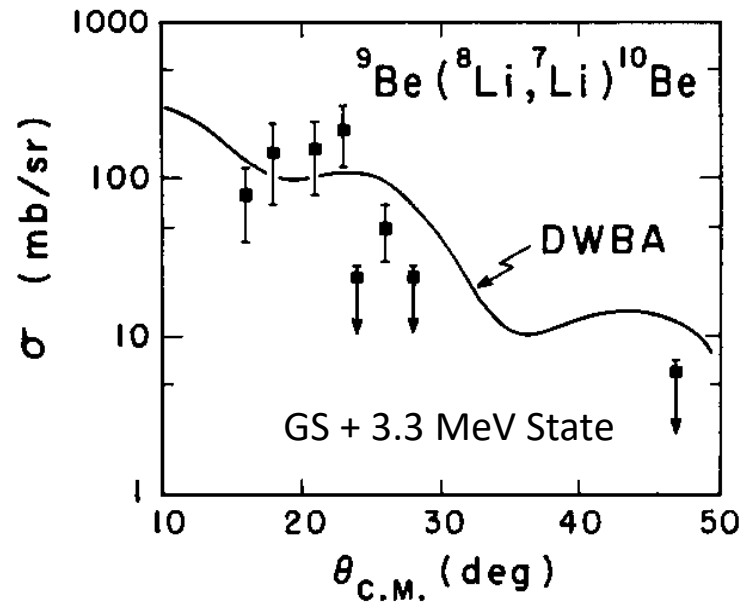
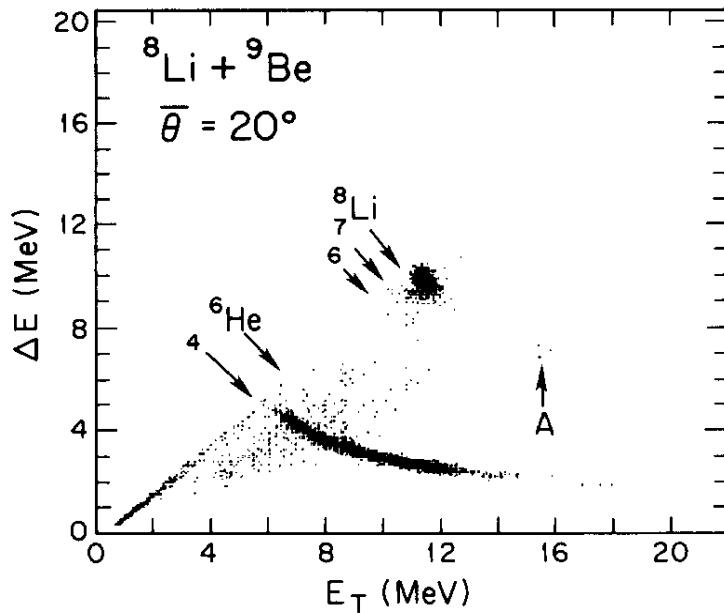
Typical Beam Parameters ( $^8\text{Li}$ ):

Primary Beam:  $^7\text{Li}$  ; 17 MeV ; 100 enA  
 $3^+$  Charge State ; Spot Size 2 mm dia.

Primary Target:  $^9\text{Be}$  ; 2.3 mg/cm<sup>2</sup>

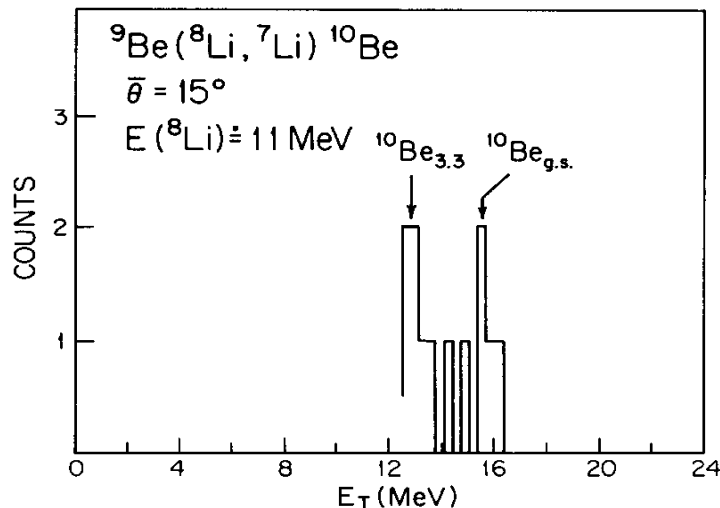
Secondary Beam:  $^8\text{Li}$  ; 14.3 MeV  
Resolution 600 keV FWHM  
Rate:  $5 \times 10^4$ /s ; Spot Size 5 mm dia.  
Divergence:  $\pm 4^\circ$

# Early Experiments: Transfer Reactions



Phys. Rev. C40, R1104 (1989)

First-Ever Successful Transfer-Reaction Measurement with a Radioactive Beam (apart from  ${}^3\text{H}$ ).



# Early Experiments: Nuclear Astrophysics

The  $^8\text{Li}(\alpha, n)^{11}\text{B}$  reaction and primordial nucleosynthesis ; *Phys. Lett. B343, 31 (1995)*

## Collaboration:

Ohio State  
Notre Dame  
Univ. of Michigan  
Kyushu Univ.  
RIKEN (Japan)  
Lawrence Berkeley Lab.  
Ball State Univ.

## Method:

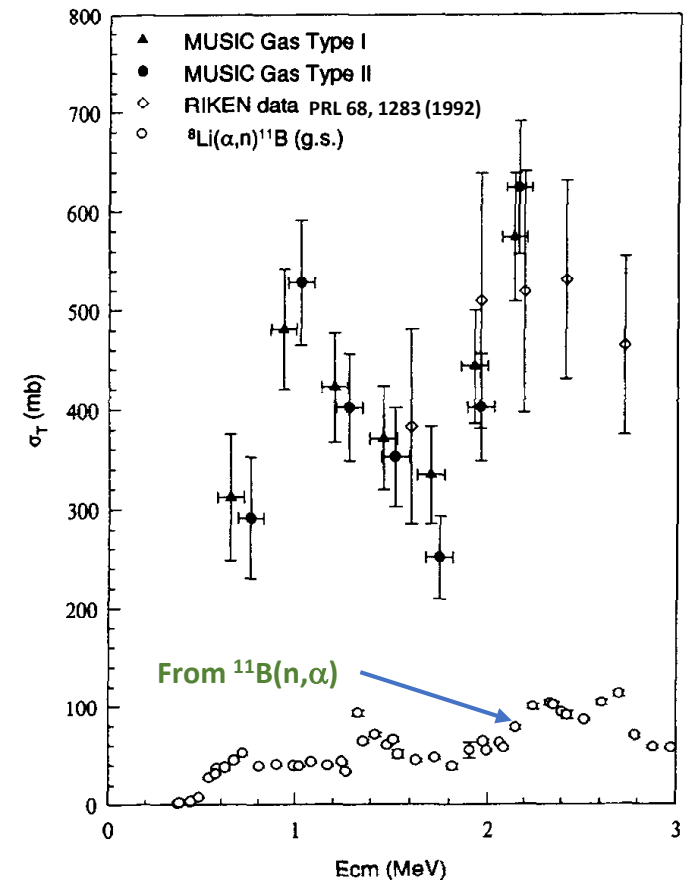
Parallel Plate  
Avalanche Counter  
(PPAC)

Multi-sampling Ion  
Chamber (MUSIC)

$^4\text{He}$  gas with 2%  
Isobutane (Type I) or  
Methane (Type II)

## Details:

$^8\text{Li}$  Energy of 7.1 MeV upon entry to the MUSIC chamber  
Energy width: 0.1 MeV ; Uncertainty in the mean: 0.2 MeV



# Early Experiments: Nuclear Astrophysics

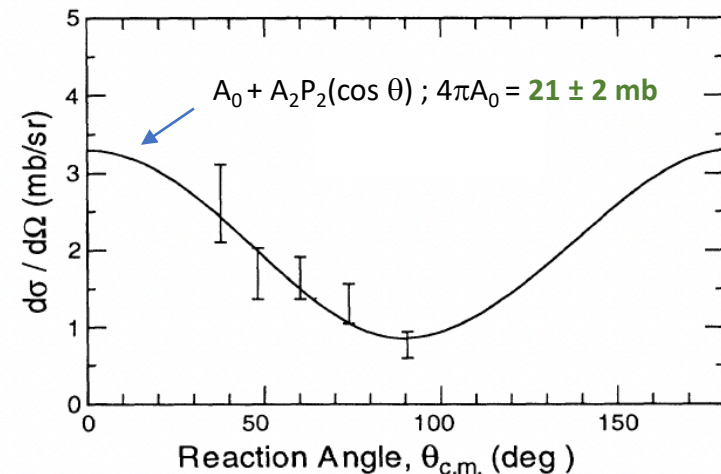
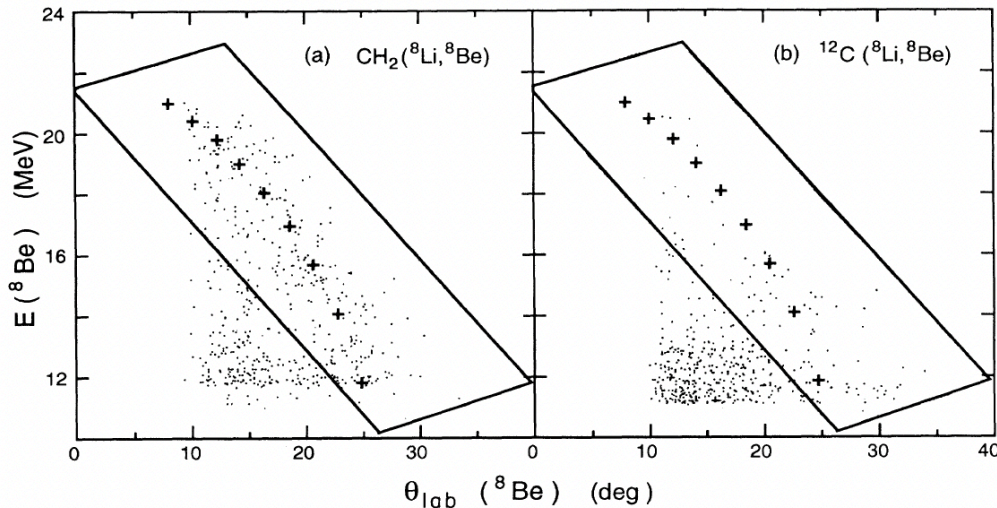
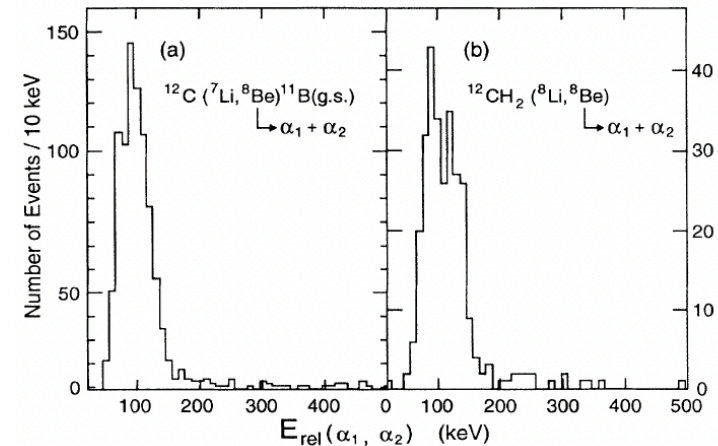
The  $^8\text{Li} (p,n)^8\text{Be}_{\text{gs}}$  reaction at  $E_{\text{cm}}=1.5$  MeV ; Phys. Rev. C47, 387 (1993)

## Collaboration:

Florida State  
Notre Dame  
Univ. of Michigan  
Univ. of North Florida

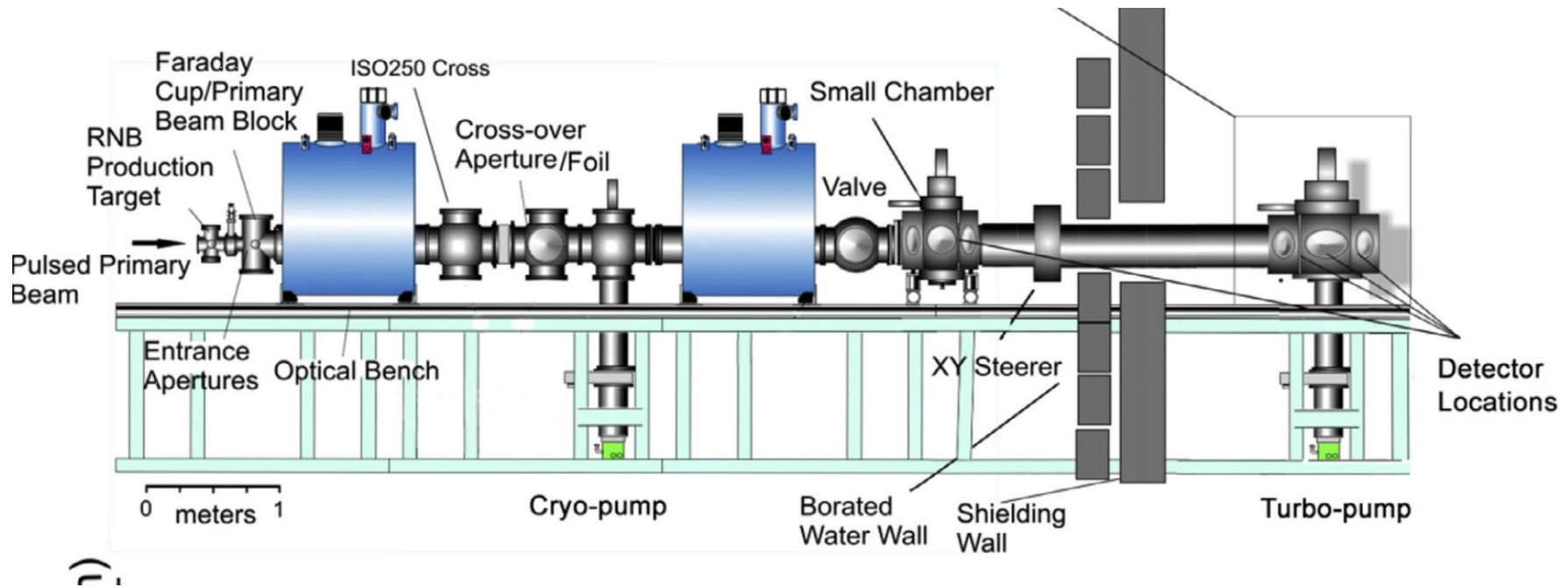
## Method:

$\alpha$ -particle coincidences  
between stacked 5 cm  
long Si strip detectors.





# New Facility: *TwinSol* (1998)



Solenoids: 6T ; 30 cm warm bore ; 1 m long

AIP Conf. Proc. 392, 397 (1997)

# Sub-Barrier Fusion with Exotic Beams

## Questions to answer:

Is fusion **suppressed** due to breakup of  ${}^6\text{He}$ ?

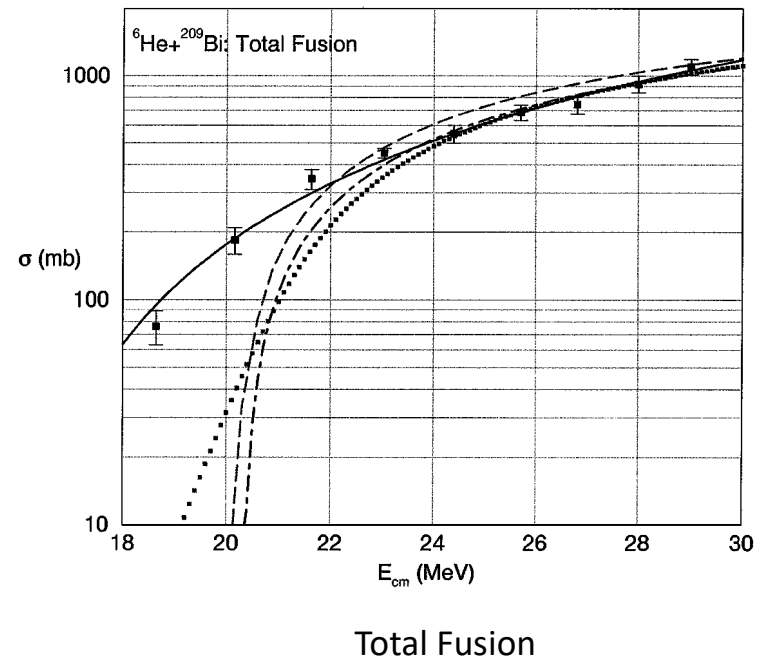
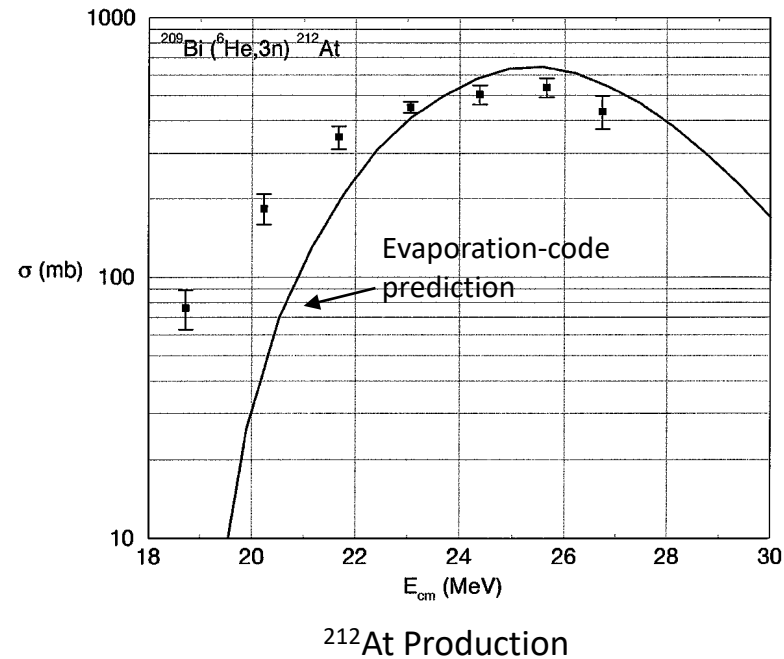
Is fusion **enhanced** due to neutrons at large radius?

## Method:

${}^6\text{He}$  incident on  ${}^{209}\text{Bi}$ .

Radiochemical: Stacked foils with degraders in-between.

Off-line detection of delayed alphas arising from evaporation residues.



Phys. Rev. Lett 81, 4580 (1998); 300 references (incl. mention in CERN Courier)



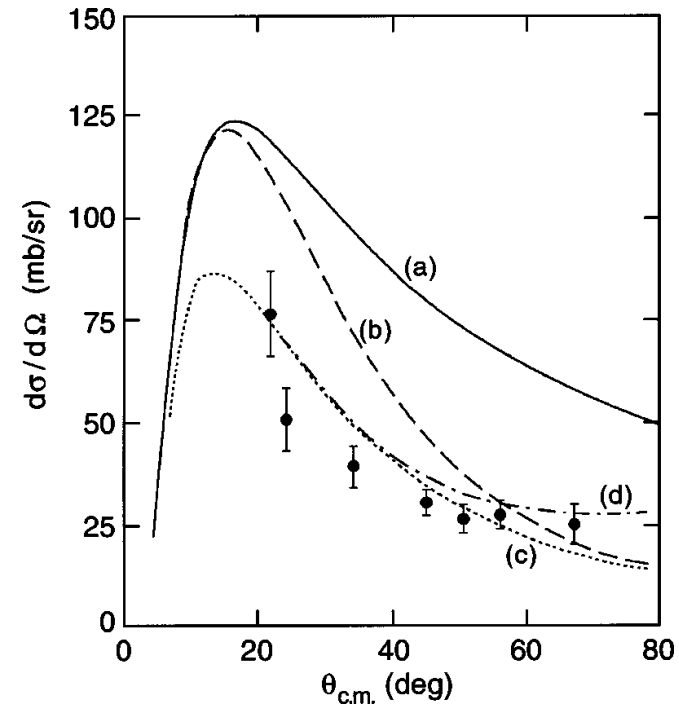
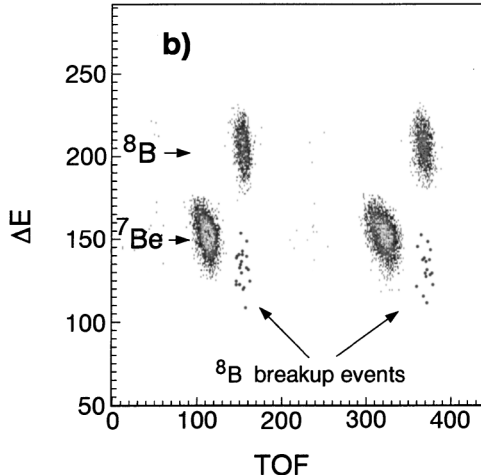
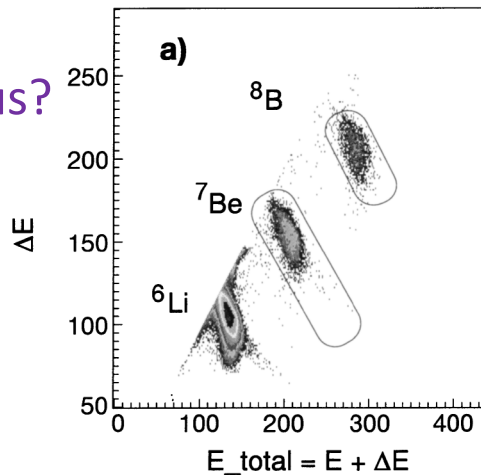
# Breakup of $^8\text{B}$ on $^{58}\text{Ni}$

Question: Is  $^8\text{B}$  a 1p halo nucleus?

Method: Add Time-of-Flight using a 2 ns FWHM pulsed beam to tag the “cocktail” beam.

## Results:

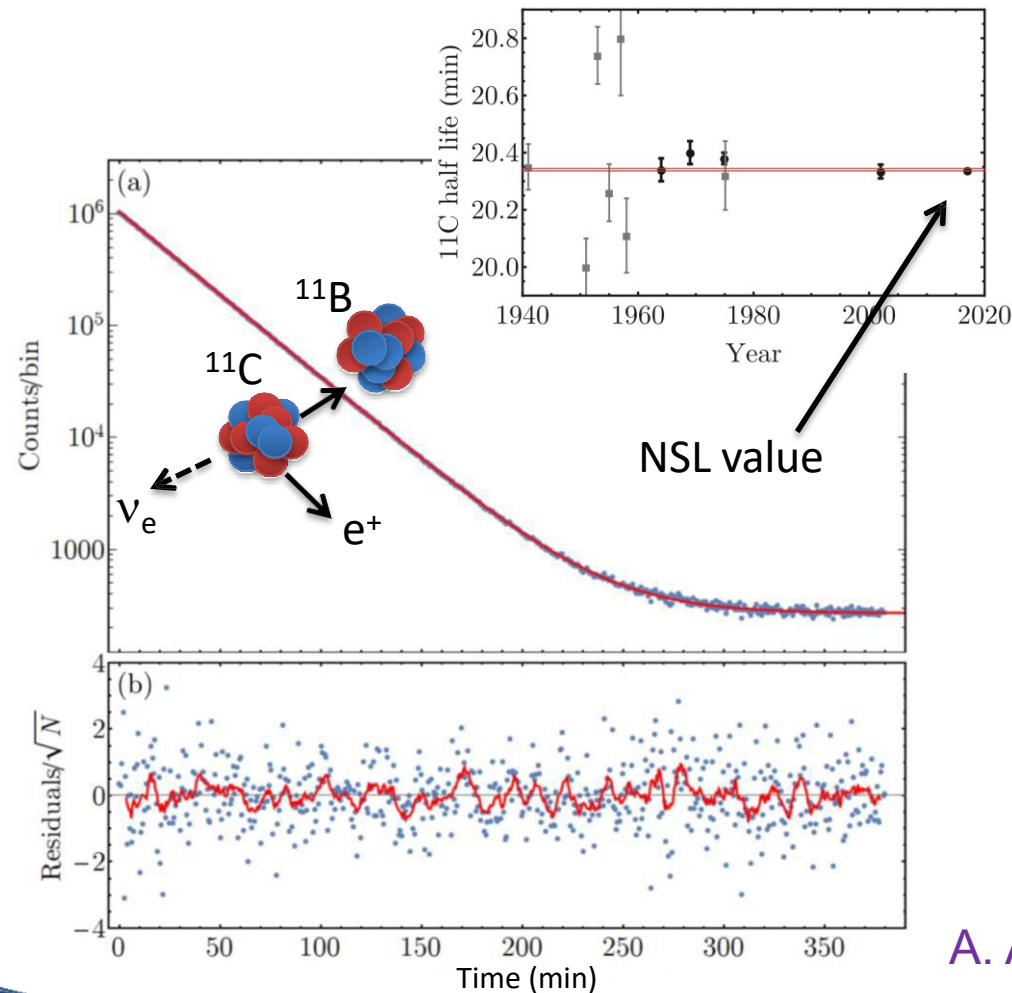
- (a) First-order Coulomb excitation assuming the usual “far-field” approximation.
- (b) Corrected for close and distant collisions due to **Coulomb polarization** of the  $^8\text{B}$  proton halo state (shielding effect).
- (c) Dynamic Coulomb dissociation of the halo state.
- (d) Nuclear effects included. Transfer process greatly suppressed due to prior breakup.



PRC66, 044609 (2002)  
Esbensen and Bertsch

PRL84, 1862 (2000)

# $^{11}\text{C}$ half-life measurement at the NSL



- Superaligned mixed mirror decays  $\rightarrow$  precision tests of Standard Model via  $V_{ud}$
- Measurement of  $^{11}\text{C}$   $t_{1/2}$  @ ND  $\rightarrow$  most precise of all SA mixed decays
- Lightest of all mirror transitions to test CVC
- Measurement of mixing ratio with the future **St. Benedict** ion trap is planned to extract  $V_{ud}$ .

A. A. Valverde *et al.*, PRC **97** 035503 (2018)

Shown by Alena Opper (NSF) at the 2019 NSAC meeting

# *TriSol* Exotic Beam Facility

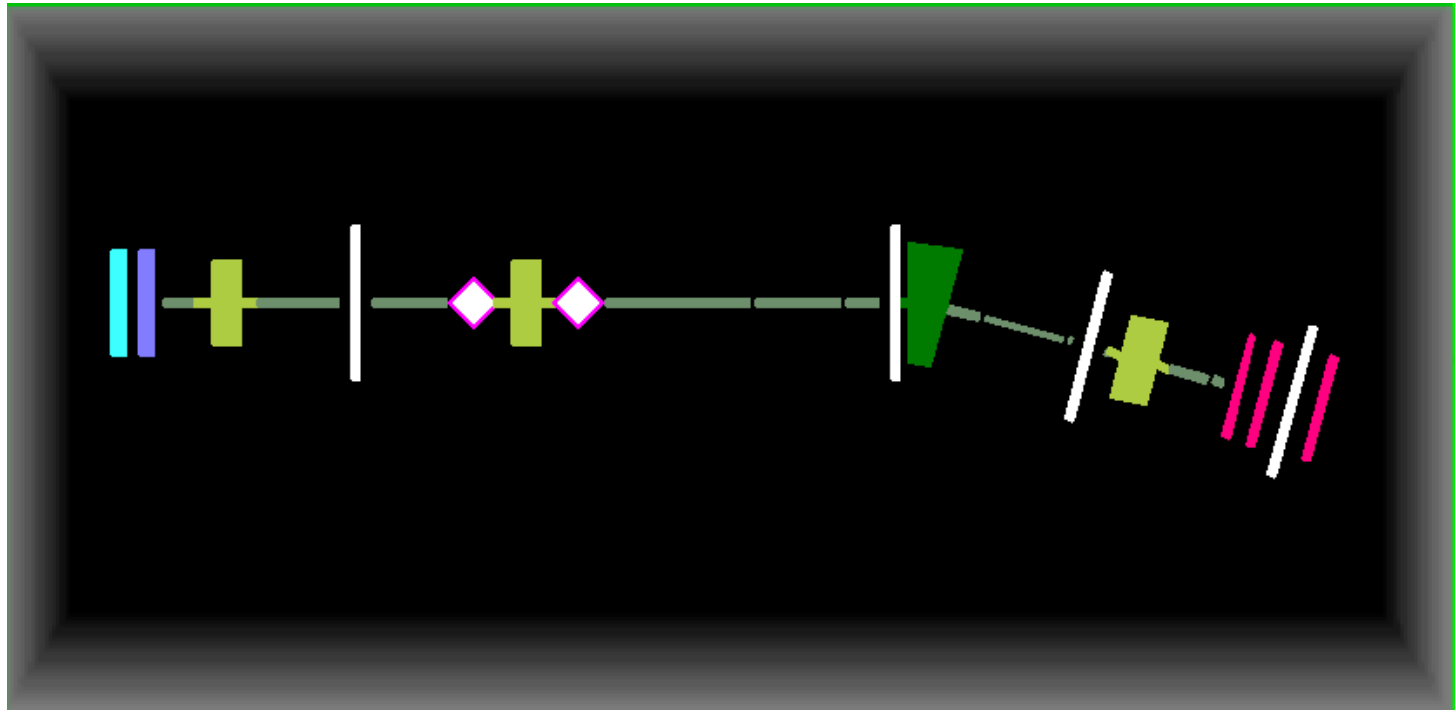


Diagram from a LISE<sup>++</sup> Calculation

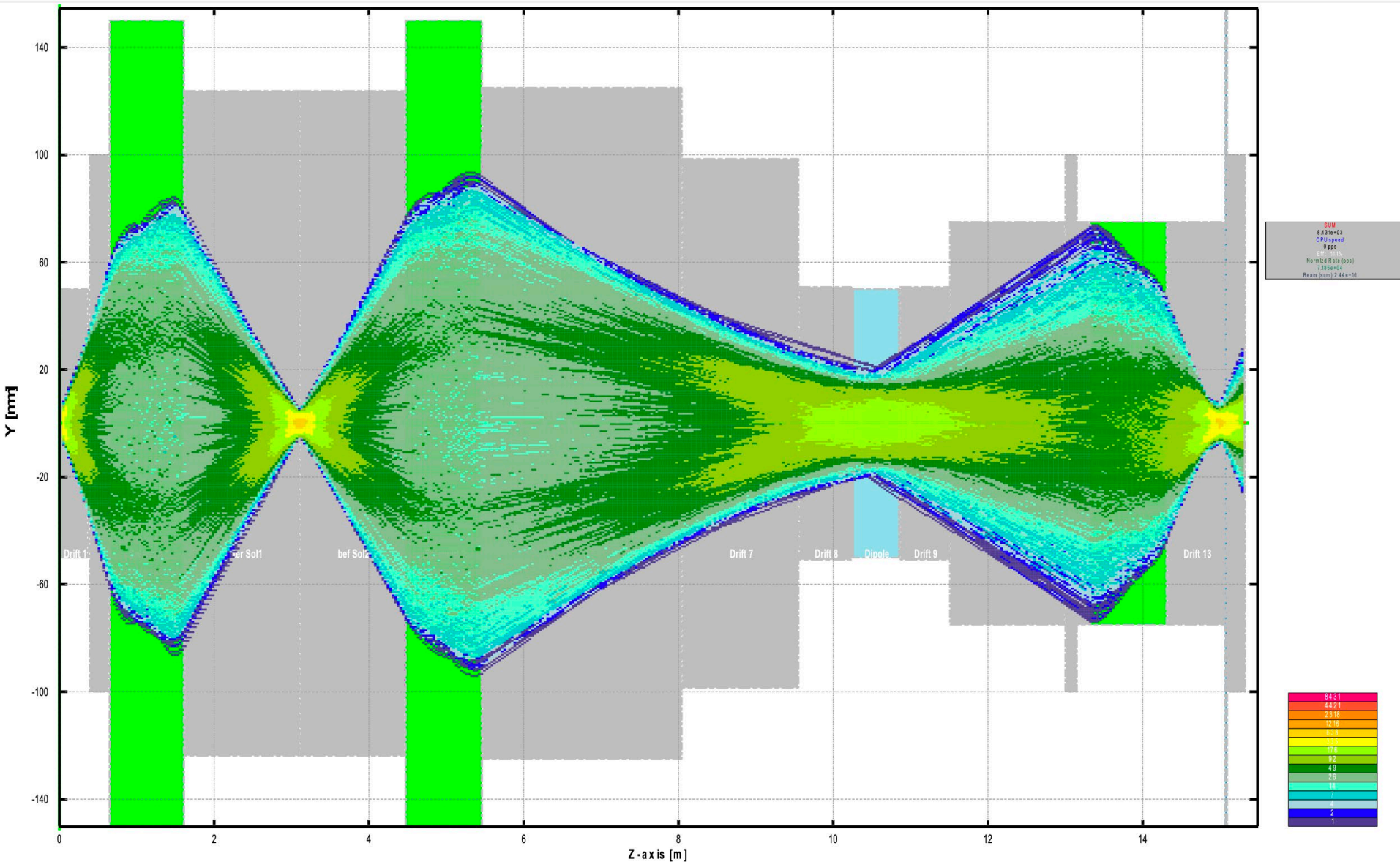
## $^8\text{B}$ : MC Transmission Plot - Envelope (only passed)

[Continue](#)

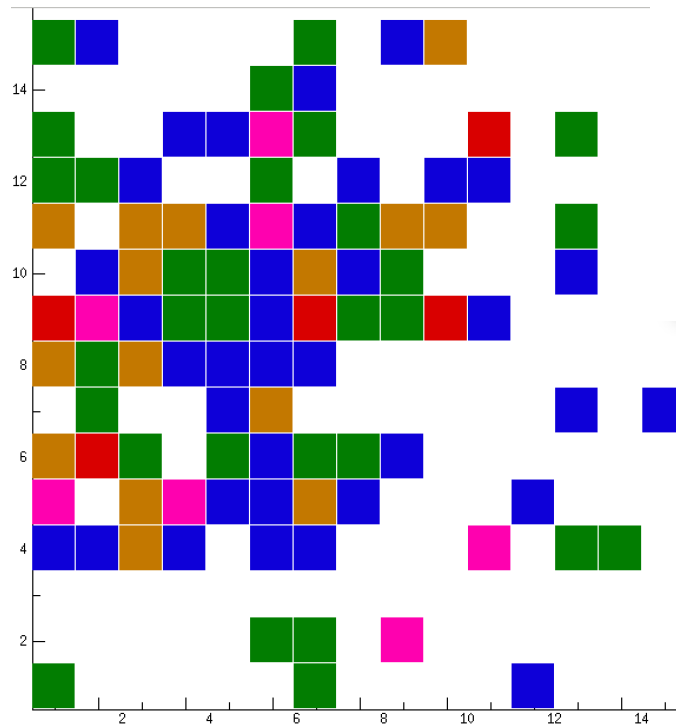
$^6\text{Li}$  (5.9 MeV/u) + He (25.4 mm), Ti (4  $\mu\text{m}$ ); Transmitted Fragment  $^8\text{B}^{5+..5+}$  (TwoBody); Optics Order: 2

dp/p=25.04%; Brho(Tm): 0.4304

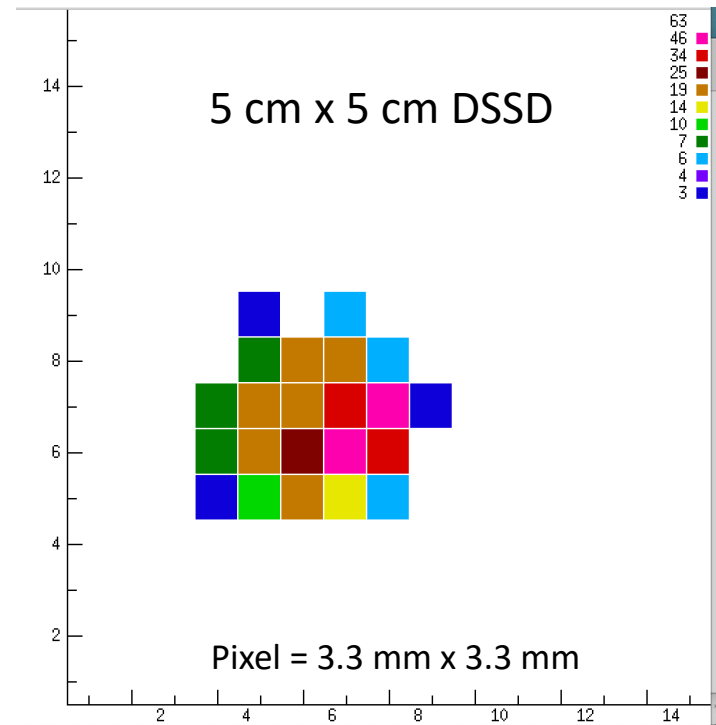
AngAccept: ON; Bounds: ON; "Drift 14" - last block for MC calc; Gate 2: "AND" (Radial [mm]); Config: dddLd|dhLhddd|Dddd|dLd\_m d\_mm|lm



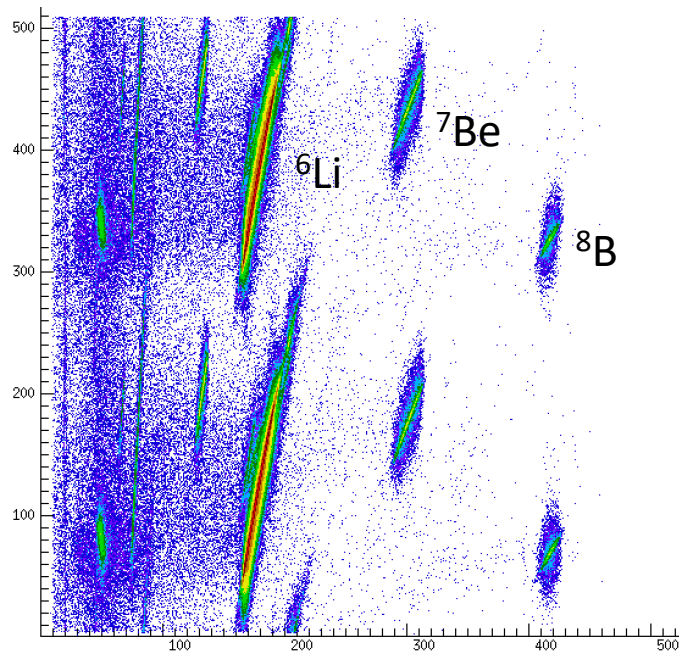
Sol3 = 0



Sol3 = 22 A



Time-of-Flight  
vs. Energy



# Small, University Based Accelerator Laboratories --- Past and Future

**Ubiquitous in the 1960's but Nearing Extinction in the 2020's**

**But**, still able to carry out interesting, important, world-class research.

Advantages: Short Turn-Around Time for Innovative Experiments.

Ability to Dedicate Long Blocks of Time for Important Work.

Hands-on Experience with All Aspects of an Experiment.

Opportunities Exist in Nuclear Astrophysics, Fundamental Interactions, and even Reactions of Exotic Beams.

**Find your niche and exploit it!**